CS-WZW CORRESPONDENCE IN OSFT

Martin Schnabl Institute of Physics, Prague Academy of Sciences of the Czech Republic

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Open String Field Theory as Fundaments of Gravity ? Closed string field theory (Zwiebach 1990) presumably THE fundamental theory of gravity — is given by a technically ingenious construction, but it is hard to work with, See i.e. N. Moeller, and also it may be viewed as too perturbative and `effective'.

 Open string field theory (Witten 1986, Berkovits 1995) might be more fundamental Sen, see also I.
Sachs and M. Baumgartl. Related to holography (Maldacena) and old 1960's ideas of Sacharov.

First look at OSFT

Open string field theory uses the following data

$$\mathcal{H}_{BCFT}, *, Q_B, \langle . \rangle.$$

Let all the string degrees of freedom be assembled in

$$|\Psi\rangle = \sum_{i} \phi_i(X) |i\rangle.$$

Witten (1986) proposed the following action

$$S = -\frac{1}{g_o^2} \left[\frac{1}{2} \left\langle \Psi * Q_B \Psi \right\rangle + \frac{1}{3} \left\langle \Psi * \Psi * \Psi \right\rangle \right],$$

First look at OSFT

This action has a huge gauge symmetry

$$\delta \Psi = Q_B \Lambda + \Psi * \Lambda - \Lambda * \Psi,$$

provide that the start product is associative, Q_B acts as graded derivation and < . > has properties of integration.

Note that there is a gauge symmetry for gauge symmetry so one expects infinite tower of ghosts – indeed they can be naturally incorporated by lifting the ghost number restriction on the string field.

Witten's star product

Defined by gluing three strings:





Pa

 P_2

 P_1

$$(\Psi_1 \star \Psi_2) \left[X(\sigma) \right] = \int \left[\mathcal{D} X_{\text{overlap}} \right] \Psi_1[\hat{X}(\sigma)] \Psi_2 \left[\check{X}(\sigma) \right]$$

It used to be a very complicated definition...



The elements of string field star algebra are states in the BCFT, they can be identified with a piece of a worldsheet.

By performing the path integral on the glued surface in two steps, one sees that in fact:

$$|\phi_1\rangle * |\phi_2\rangle = |\phi_1 e^{-K} \phi_2\rangle.$$

Witten's star product as operator multiplication

We have just seen that the star product obeys

$$|\phi_1\rangle * |\phi_2\rangle = |\phi_1 e^{-K} \phi_2\rangle.$$

And therefore states $\hat{\phi} = e^{K/2} \phi e^{K/2}$ obey

$$|\hat{\phi}_1\rangle * |\hat{\phi}_2\rangle = |\widehat{\phi_1\phi_2}\rangle$$

The star product and operator multiplication are thus isomorphic!

In case you wonder what e^{-K} is

To find out what e^{-K} stands for, one can perform a conformal transformation





 $e^{-K} = e^{-rac{\pi}{2}\int_{-M}^{M}T_{ ilde{z} ilde{z}}d ilde{z}}$ Just like $e^{-tL_0} = e^{-t\oint T_{ww}dw}$

Simple subsector of the star algebra

- The star algebra is formed by vertex operators and the operator *K*. The simplest subalgebra relevant for tachyon condensation is therefore spanned by *K* and *c*. Let us be more generous and add an operator *B* such that *QB*=*K*.
- The building elements thus obey

$$c^2 = 0, \quad B^2 = 0, \quad \{c, B\} = 1$$

 $[K, B] = 0, \quad [K, c] = \partial c$

■ The derivative *Q* acts as

$$Q_B K = 0, \quad Q_B B = K, \quad Q_B c = c K c.$$

Classical solutions

This new understanding lets us construct solutions to OSFT equations of motion $Q_B\Psi + \Psi * \Psi = 0$ easily.

It does not take much trying to find the simplest solution is $\Psi = \alpha c - cK$ $Q\Psi = \alpha(cKc) - (cKc)K$ $\Psi * \Psi = \alpha^2 c^2 - \alpha c^2 K - \alpha cKc + (cK)(cK)$

More general solutions are of the form

$$\Psi = Fc \frac{KB}{1 - F^2} cF,$$

Here F=F(K) is arbitrary Schnabl 2005, Okawa, Erler 2006

- The space of all such solutions has not been completely classified although we are close quite close (Erler).
- Let us restrict our attention to different choices of F(K) only.

• Let us call a state *geometric* if F(K) is of the form

$$F(K) = \int_0^\infty d\alpha f(\alpha) e^{-\alpha K}$$

where $f(\alpha)$ is a tempered distribution

Therefore F(K) must be holomorphic for Re(K)>0 and bounded by a polynomial there.

• Since formally $\Psi = (1 - FBcF)Q(1 - FBcF)^{-1}$,

and
$$(1 - FBcF)^{-1} = 1 + \frac{F}{1 - F^2}BcF$$

the state is trivial if $F/(1 - F^2)$ is well defined

There is another useful criterion. One can look at the cohomology of the theory around a given solution. It is given by an operator

$$Q_{\Psi} = Q_B + \{\Psi, \cdot\}_*.$$

• The cohomology is formally trivialized by an operator $A = \frac{1 - F^2}{K} B$, which obeys $\{Q_{\Psi}, A\} = 1$.

- Therefore in this in this class of solutions, the trivial ones are those for which $F^2(0) \neq 1$.
- Tachyon vacuum solutions are those for which $F^2(0) = 1$ but the zero of $1 F^2$ is first order
- When the order of zero of 1-F² at K=0 is of higher order the solution is not quite well defined, but it has been conjectured (Ellwood, M.S.) to correspond to multibrane solutions.

Examples

 $F(K) = a \quad (const.)$ $F(K) = \sqrt{1 - \beta K}$ $F(K) = e^{-K}$ $F(K) = \frac{1}{\sqrt{1 + K}}$

... trivial solution

.... 'tachyon vacuum' only *c* and *K* turned on

.... M.S. '05

... Erler, M.S. '09 – the simplest solution so far

Towards the super OSFT

- There has been much progress on many fronts in our field, we are finding new solutions (new classical backgrounds), study cohomology, couplings to closed strings etc.
- One of the outstanding problems is to find a generalization to open superstrings. One of the most popular actions is due to Berkovits:

$$S = \frac{1}{2g^2} \left\langle \left(e^{-\Phi} Q e^{\Phi} \right) \left(e^{-\Phi} \eta_0 e^{\Phi} \right) - \int_0^1 dt \left(e^{-t\Phi} \partial_t e^{t\Phi} \right) \left(e^{-t\Phi} Q e^{t\Phi} \right) \left(e^{-t\Phi} \eta_0 e^{t\Phi} \right) \right\rangle,$$

since it is non-polynomial, it has been very hard to work with.

CS-WZW correspondence

In our hopefully-soon-to-appear work with P.A. Grassi we borrowed some classical results from the CS-WZW correspondence and rewrote the action as

$$S = -\frac{1}{g^2} \left[\frac{1}{2} \left\langle \! \left\langle \mathbb{A} \ast \mathbb{Q} \mathbb{A} \right\rangle \! \right\rangle + \frac{1}{3} \left\langle \! \left\langle \mathbb{A} \ast \mathbb{A} \ast \mathbb{A} \right\rangle \! \right\rangle \right].$$

 In principle, in this form the action should be amenable to straightforward quantization (work under progress).

CS-WZW correspondence

 Let us consider the following 2+1 dimensional theory with gauge group G defined on a manifold Σ×R

$$S_{CS} = -\frac{1}{2g^2} \int_{\Sigma \times \mathbb{R}} \operatorname{Tr}(AdA + \frac{2}{3}A^3),$$

□ Let $A_0=0$ on the boundary $\partial \Sigma$. Let us decompose $A=A_0+\tilde{A}$. Then

$$S = -\frac{k}{4\pi} \int_{\mathbf{Y}} \operatorname{Tr}\left(\tilde{A}\frac{\partial}{\partial t}\tilde{A}\,\mathrm{d}t\right) + \frac{k}{2\pi} \int_{\mathbf{Y}} \operatorname{Tr}\left[A_0(\tilde{\mathbf{d}}\tilde{A} + \tilde{A}^2)\right]$$

CS-WZW correspondence

 After integrating out A₀ we get δ(F'), where
F'=d' Ã +Ã Ã. This constraint can be easily solved by Ã=U d' U⁻¹

• Let us use U as the integration variable instead of \tilde{A}

$$\int \mathrm{D}\tilde{A}\,\delta(\tilde{F}) = \int \mathrm{D}U$$

The Jacobian of this transformation is one

$$S = kS_{\rm C}^+(U) \equiv \frac{k}{4\pi} \int_{\partial Y} \operatorname{Tr}\left(U^{-1}\partial_{\phi}UU^{-1}\partial_{t}U\right) \mathrm{d}\phi \,\mathrm{d}t + \frac{k}{12\pi} \int_{Y} \operatorname{Tr}\left(U^{-1}\mathrm{d}U\right)^3$$

Elitzur, Moore, Schwimmer, Seiberg 1989

CS-WZW correspondence for the super OSFT

□ Let us tensor the large Hilbert space of super OSFT with space of differential forms on the unit interval parameterized by t ∈ [0, 1],
□ This makes all elements of the Hilbert space *t*-dependent, and also doubles it

• Let us assign ghost number one to dt, and let $\mathbb{Q} = Q + \eta_0 + d_t$, where $d_t = dt \frac{\partial}{\partial t}$.

Note that this operator has the correct cohomology as $H(Q + \eta_0 + d_t) \simeq H(Q + \eta_0, H(d_t))$.

CS-WZW correspondence for the super OSFT

• Let us further define $\langle \langle \dots \rangle \rangle = \int_0^1 \langle \dots \rangle$ so that

$$\langle\!\langle \mathbb{A}_1 * \mathbb{A}_2 * \ldots * \mathbb{A}_n \rangle\!\rangle = \langle\!\langle \mathbb{A}_2 * \ldots * \mathbb{A}_n * \mathbb{A}_1 \rangle\!\rangle.$$

and

 $\langle\!\langle Q \mathbb{X}(t) \rangle\!\rangle = 0, \quad \langle\!\langle \eta_0 \mathbb{X}(t) \rangle\!\rangle = 0, \quad \langle\!\langle d_t \mathbb{X}(t) \rangle\!\rangle = \langle \mathbb{X}(1) - \mathbb{X}(0) \rangle$

CS-WZW correspondence for the super OSFT

The proposed action takes the form

$$S = -\frac{1}{g^2} \left[\frac{1}{2} \left\langle \! \left\langle \mathbb{A} \ast \mathbb{Q} \mathbb{A} \right\rangle \! \right\rangle + \frac{1}{3} \left\langle \! \left\langle \mathbb{A} \ast \mathbb{A} \ast \mathbb{A} \right\rangle \! \right\rangle \right].$$

where at the classical level the string field can be restricted to

$$\mathbb{A} = A^{(1,0,0)} + B^{(1,-1,0)} + C^{(0,0,1)},$$

What next?

Complete the construction at the quantum level using the BV formalism, it is very nontrivial to solve the master equation (work in progress with Berkovits, Okawa, Torii, Zwiebach and Grassi)

- Find the tachyon vacuum solutions on non-BPS branes.
- Study brane-antibrane tachyon driven dynamics (and apply to early universe models)
- When you are done, go back and look for closed strings.