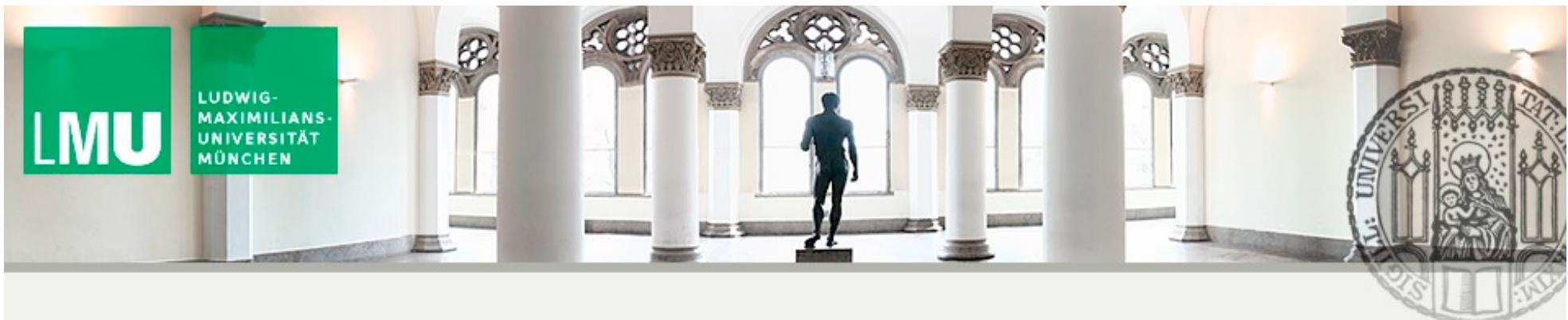


# String Amplitudes for the LHC in D-brane Compactifications

Dieter Lüst, LMU (Arnold Sommerfeld Center)  
and MPI München



# Introduction:

Count the number of consistent string vacua ➤

Vast landscape with  $N_{sol} = 10^{500-1500}$  vacua!

(Kawai, Lewellen, Tye (1986); Lerche, Lüst, Schellekens (1986);  
Antoniadis, Bachas, Kounnas (1986); Douglas (2003))

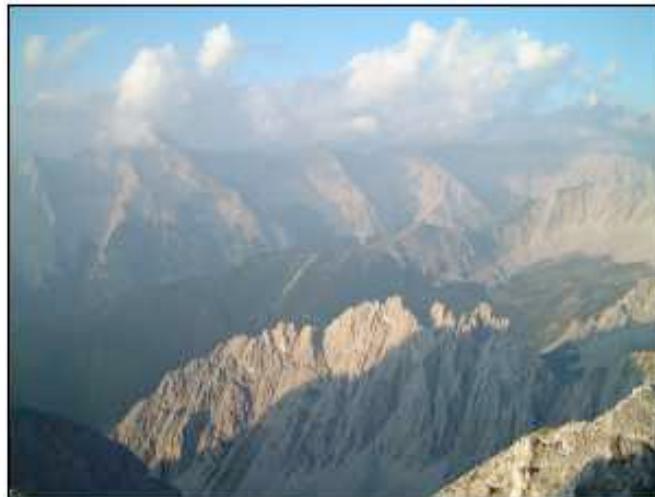


# Introduction:

Count the number of consistent string vacua ➤

Vast landscape with  $N_{sol} = 10^{500-1500}$  vacua!

(Kawai, Lewellen, Tye (1986); Lerche, Lüst, Schellekens (1986);  
Antoniadis, Bachas, Kounnas (1986); Douglas (2003))



Two strategies to find something interesting:

# Introduction:

Count the number of consistent string vacua ➤

Vast landscape with  $N_{sol} = 10^{500-1500}$  vacua!

(Kawai, Lewellen, Tye (1986); Lerche, Lüst, Schellekens (1986);  
Antoniadis, Bachas, Kounnas (1986); Douglas (2003))



Two strategies to find something interesting:

- Explore all mathematically consistent possibilities:  
**top down approach** (quite hard), string statistics.

# Introduction:

Count the number of consistent string vacua ➤

Vast landscape with  $N_{sol} = 10^{500-1500}$  vacua!

(Kawai, Lewellen, Tye (1986); Lerche, Lüst, Schellekens (1986);  
Antoniadis, Bachas, Kounnas (1986); Douglas (2003))



Two strategies to find something interesting:

- Explore all mathematically consistent possibilities:  
**top down approach** (quite hard), string statistics.
- Do not look randomly - look for green (promising) spots  
in the landscape   ⇒ model building, **bottom up approach**.

# Anthropic principle?

Our universe is not special!

(Susskind, 2003; see also  
Schellekens, arXiv:0807.3249)

Observed parameters take their observed values for the simple reason that they allow for intelligent life.

- Fine structure and strong coupling constants:  
nucleo-synthesis

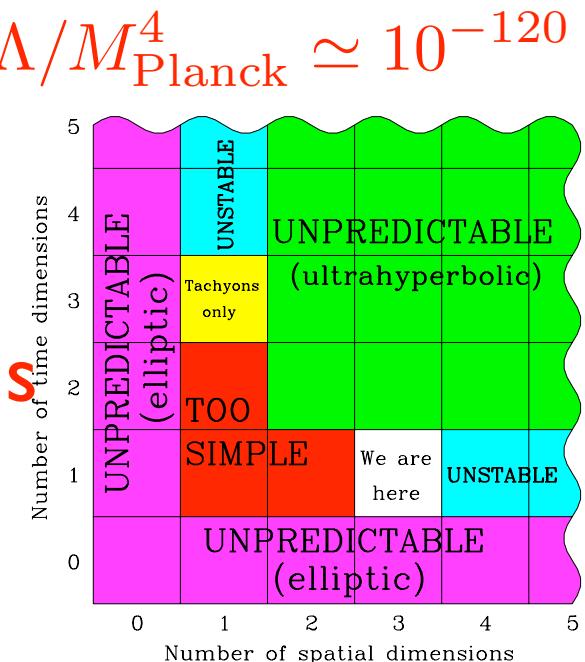
Gravity:

- Fine tuning of cosmological constant:

(Weinberg, 1987)

⇒ Need at least  $10^{120}$  vacua!

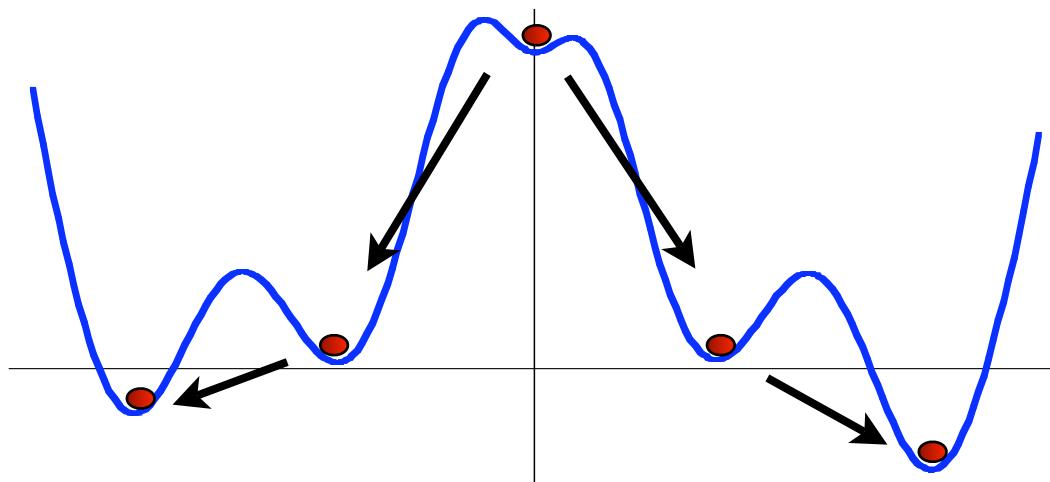
- Ehrenfest: number of spatial dimensions



# Multiverse picture: (Linde, 1986)

Transition amplitudes between different vacua (wave function of the universe):

(Hartle, Hawking, 1983)

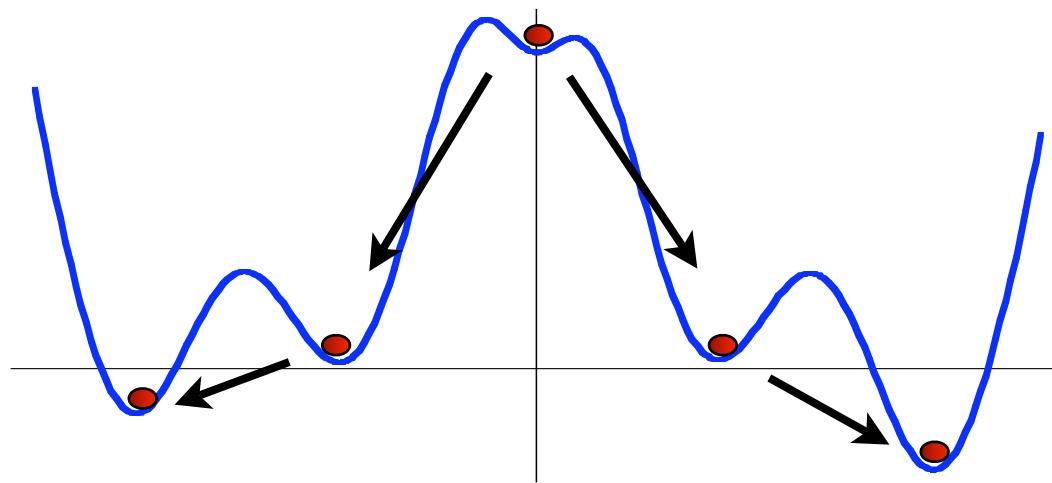


# Multiverse picture:

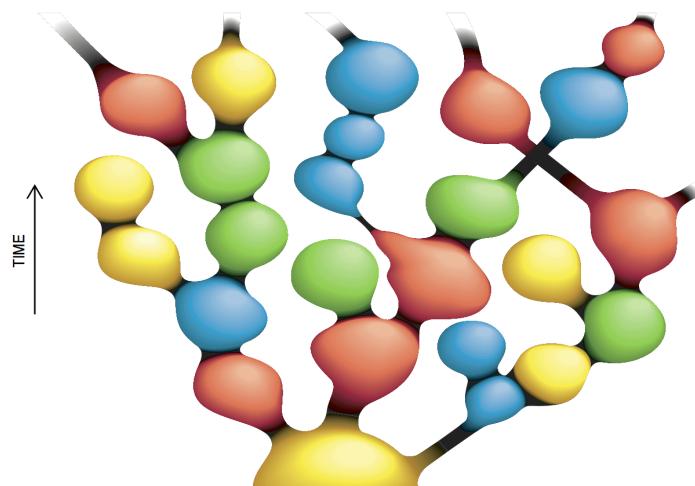
(Linde, 1986)

Transition amplitudes between different vacua (wave function of the universe):

(Hartle, Hawking, 1983)



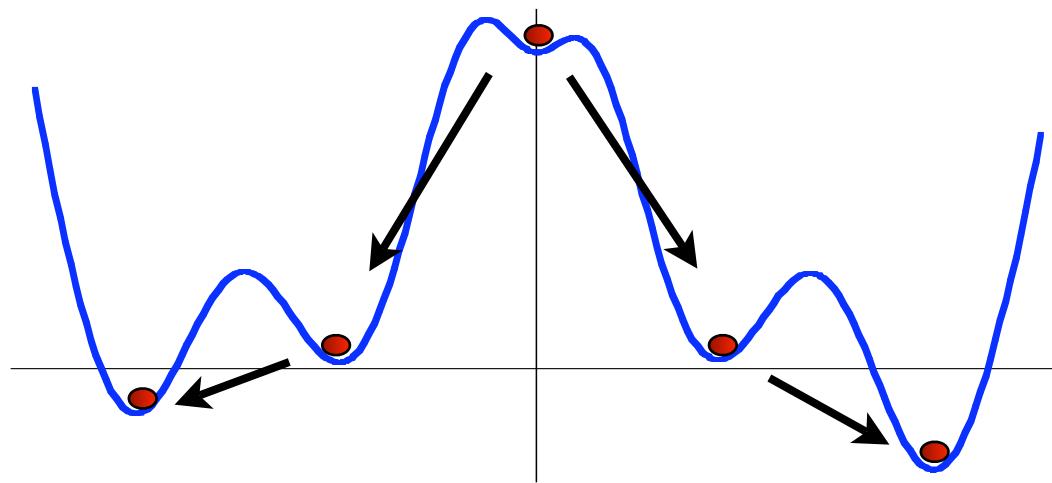
⇒ Eternal, self-producing universe:



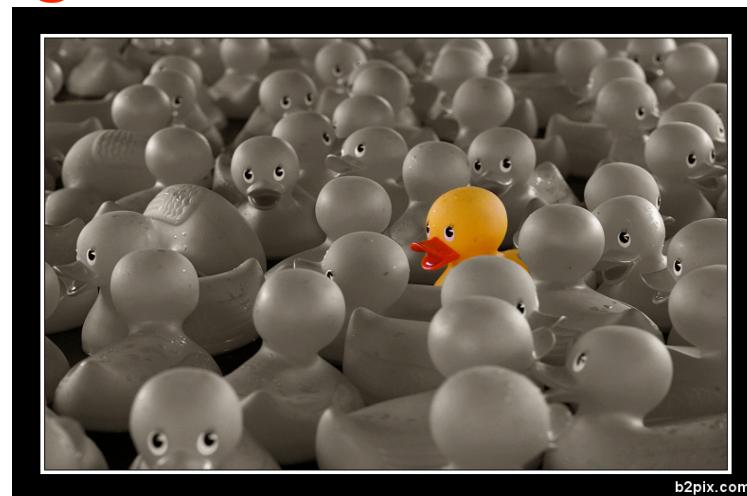
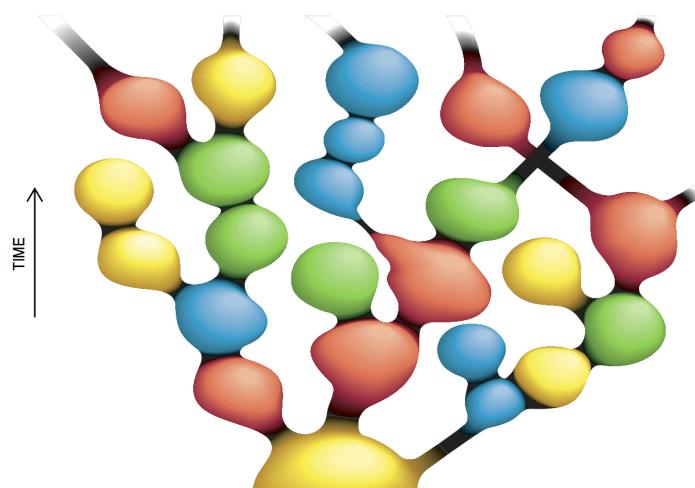
# Multiverse picture: (Linde, 1986)

Transition amplitudes between different vacua (wave function of the universe):

(Hartle, Hawking, 1983)



⇒ Eternal, self-producing universe:



b2pix.com

# **General feature of string theory: Geometrization of particles and their interactions!**

# General feature of string theory: Geometrization of particles and their interactions!

## Dictionary:

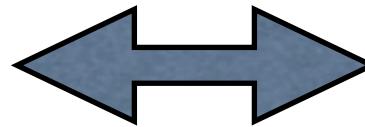
### Particles physics

Leptons	Quarks		
	u up	c charm	t top
d down	s strange	b bottom	
$\nu_e$ e- Neutrino	$\nu_\mu$ $\mu$ - Neutrino	$\nu_\tau$ $\tau$ - Neutrino	
e electron	$\mu$ muon	$\tau$ tau	
I	II	III	

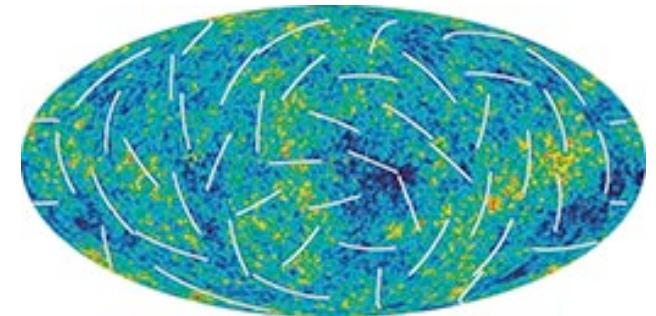
The Generations of Matter

Gauge interactions:

$$G = SU(3) \times SU(2) \times U(1)$$



### Cosmology



# General feature of string theory: Geometrization of particles and their interactions!

## Dictionary:

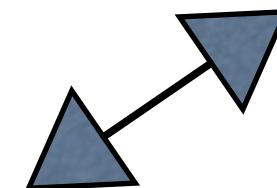
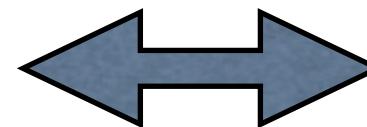
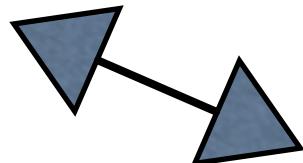
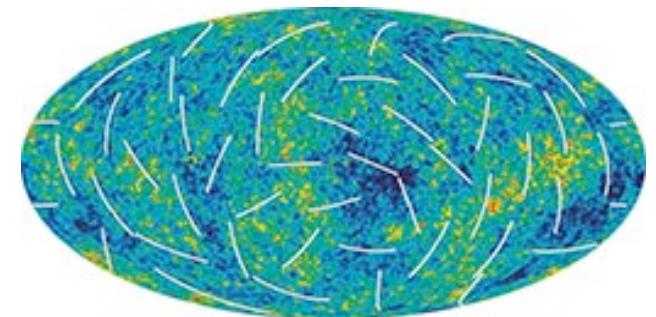
### Particles physics

Leptons	Quarks		
	u up	c charm	t top
d down	s strange	b bottom	
$\nu_e$ e- Neutrino	$\nu_\mu$ $\mu$ - Neutrino	$\nu_\tau$ $\tau$ - Neutrino	
e electron	$\mu$ muon	$\tau$ tau	
I	II	III	The Generations of Matter

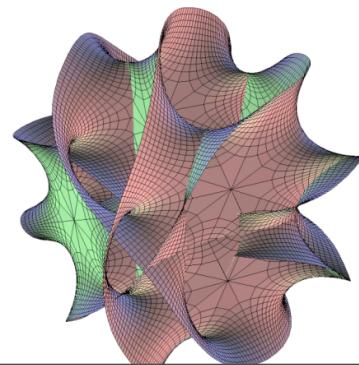
Gauge interactions:

$$G = SU(3) \times SU(2) \times U(1)$$

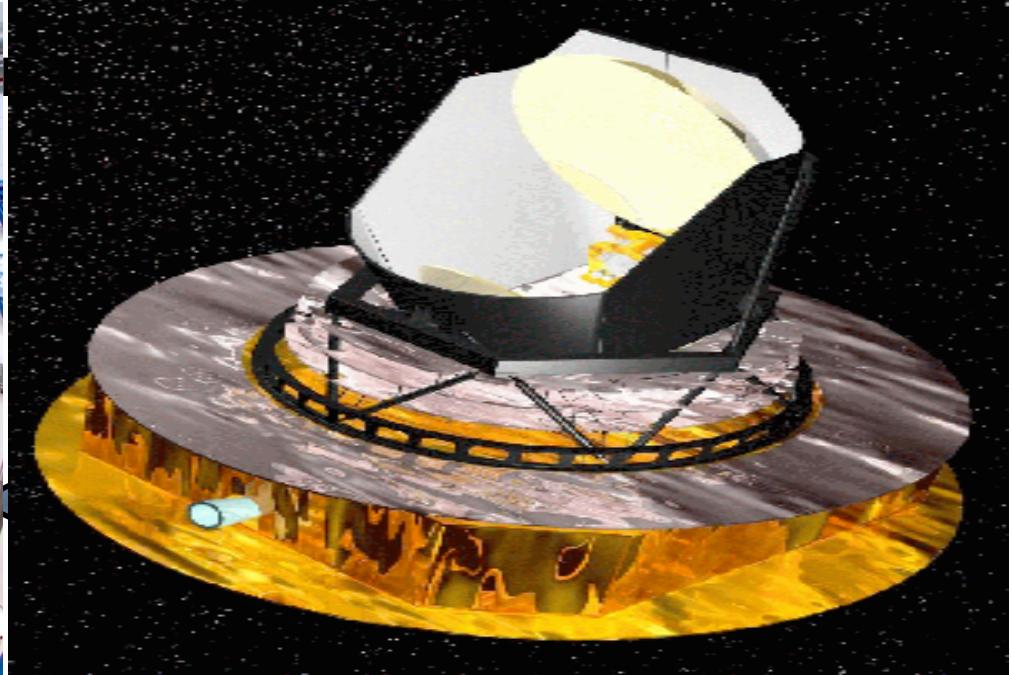
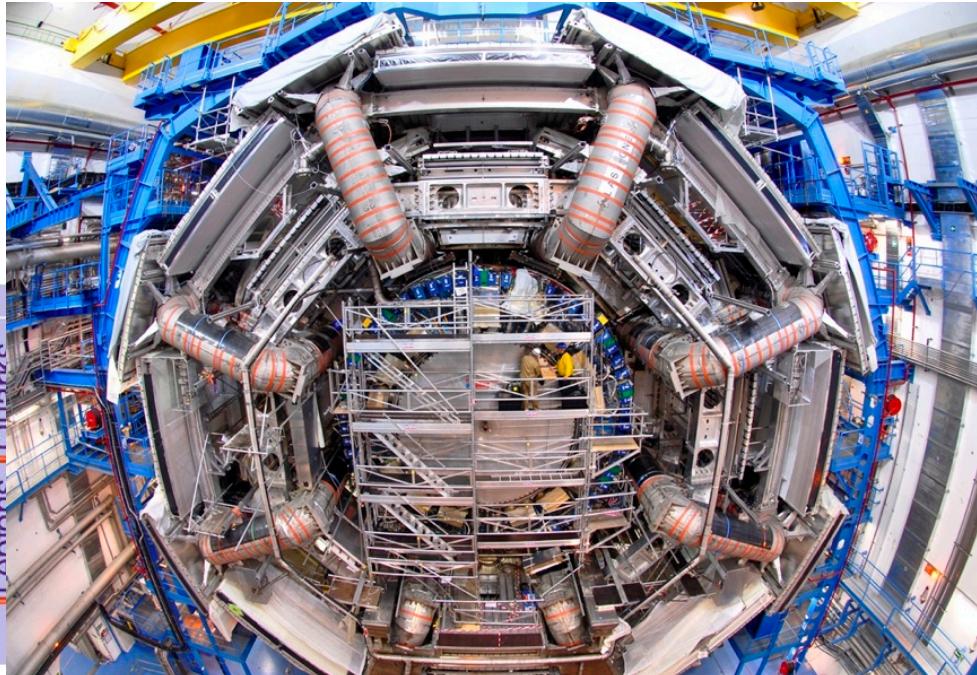
### Cosmology



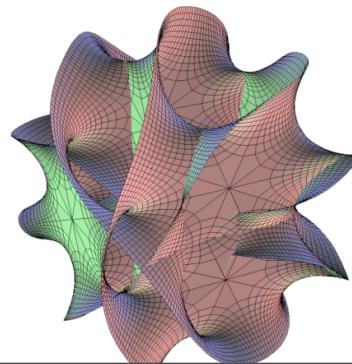
geometry & topology of strings and branes



# General feature of string theory: Geometrization of particles and their interactions!



geometry & topology of strings and branes



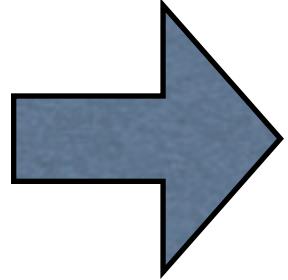
# Strategy for string phenomenology:

Consider (only) those vacua that realize the Standard Model  
(by-pass the landscape problem):

- What is the likelihood for vacua with the SM-like properties?
- What are their generic, model independent features?
- Can we make model independent predictions beyond the SM?
- Can we test these predictions in experiments (LHC)?

Bottum-up approach has to meet  
top-down approach!

# Outline

- 
- Intersecting D-brane models
  - Mass scales in D-brane models
  - Stringy amplitudes for the LHC

(The LHC string hunter's companion)

## II) (Intersecting) D-brane models:

(Bachas (1995); Blumenhagen, Görlich, Körs, Lüst (2000); Angelantonj, Antoniadis, Dudas Sagnotti (2000); Ibanez, Marchesano, Rabadan (2001); Cvetic, Shiu, Uranga (2001); ...)

### Alternative constructions: heterotic strings

(Braun, He, Ovrut, Pantev; Bouchard, Donagi; Buchmüller, Hamaguchi, Lebedev, Nilles, Ramos-Sánchez, Ratz, Vaudrevange; Groot Nibbelink, Held, Ruehle, Trapletti, Vaudrevange; Faraggi, Kounnas, Rizos)

F-theory (Beasley, Heckman, Marsano, Saulina, Schäfer-Nameki, Vafa; Donagi, Wijnholt, ...)

Consider open string compactifications with intersecting D-branes  $\rightarrow$  Type IIA/B orientifolds:

Features:

- Non-Abelian gauge bosons live as open strings on lower dimensional world volumes  $\pi$  of D-branes.
- Chiral fermions are open strings on the intersection locus of two D-branes:  $N_F = I_{ab} \equiv \#(\pi_a \cap \pi_b) \equiv \pi_a \circ \pi_b$

# Perturbative type II orientifolds contain:

(Review: Blumenhagen, Körs, Lüst, Stieberger, hep-th/0610327)

- Closed string 6-dimensional background geometry:
  - Torus, orbifold, Calabi-Yau space, generalized spaces with torsion.
- Space-time filling D(3+p)-branes wrapped around internal p-cycles:
  - Open string matter fields.
- Strong consistency conditions:
  - tadpole cancellation with orientifold planes.

D6 wrapped on 3-cycles  $\pi_a$ , intersect at angles  $\theta_{ab}$

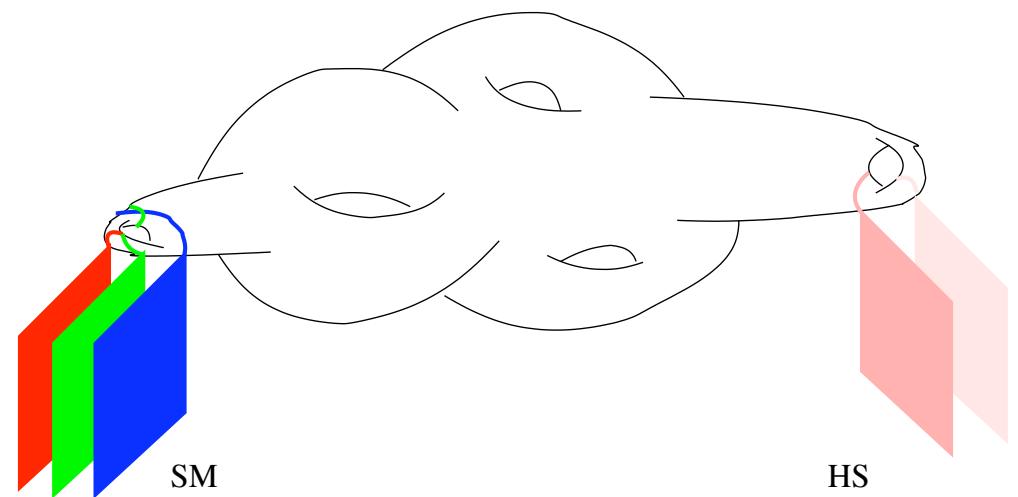
Tadpole condition:

$$\sum_a N_a \pi_a = \pi_{O6}$$

D6 wrapped on 3-cycles  $\pi_a$ , intersect at angles  $\theta_{ab}$

Tadpole condition:

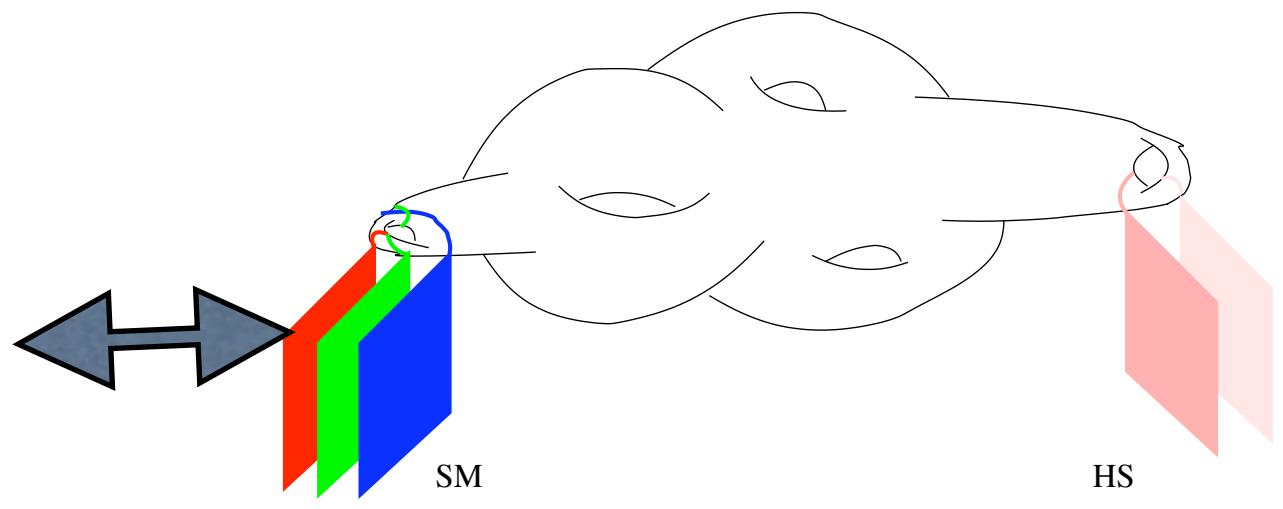
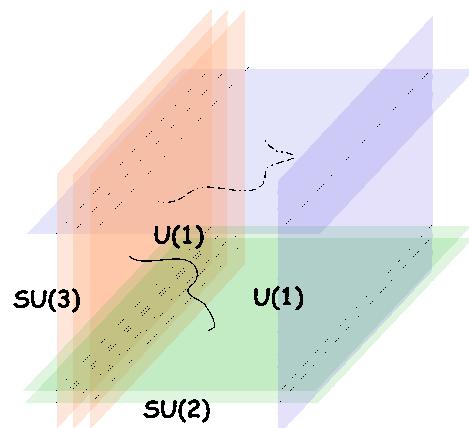
$$\sum_a N_a \pi_a = \pi_{O6}$$



D6 wrapped on 3-cycles  $\pi_a$ , intersect at angles  $\theta_{ab}$

Tadpole condition:

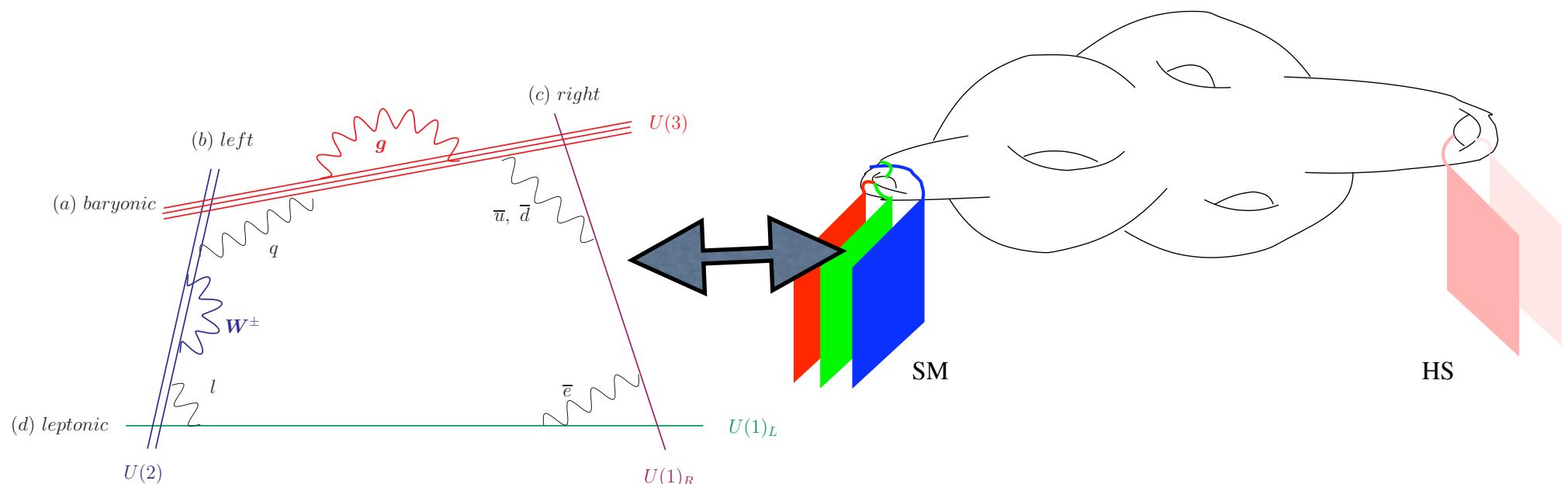
$$\sum_a N_a \pi_a = \pi_{O6}$$



D6 wrapped on 3-cycles  $\pi_a$ , intersect at angles  $\theta_{ab}$

Tadpole condition:

$$\sum_a N_a \pi_a = \pi_{O6}$$



(Ibanez, Marchesano, Rabadan, hep-th/0105155;  
Blumenhagen, Körs, Lüst, Ott, hep-th/0107138)



How many orientifold models exist which come close to the (spectrum of the) MSSM?

(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:



How many orientifold models exist which come close to the (spectrum of the) MSSM?

(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:

Z6'-orientifold: (Gmeiner, Honecker, arXiv:0806.3039)

Millions of standard models!

How many orientifold models exist which come close to the (spectrum of the) MSSM?

(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

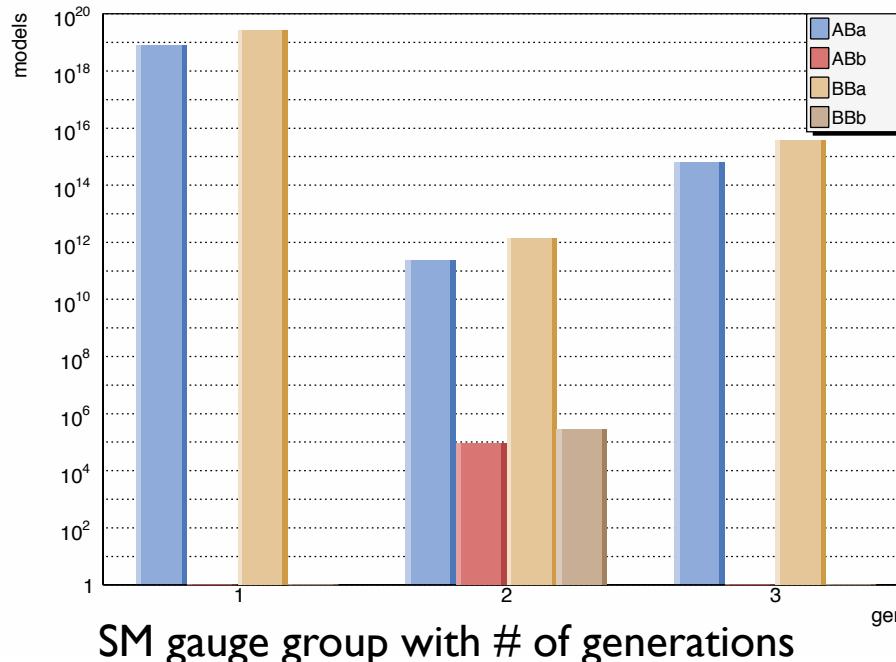
Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:

Z6'-orien

Mi



3039)



How many orientifold models exist which come close to the (spectrum of the) MSSM?

(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

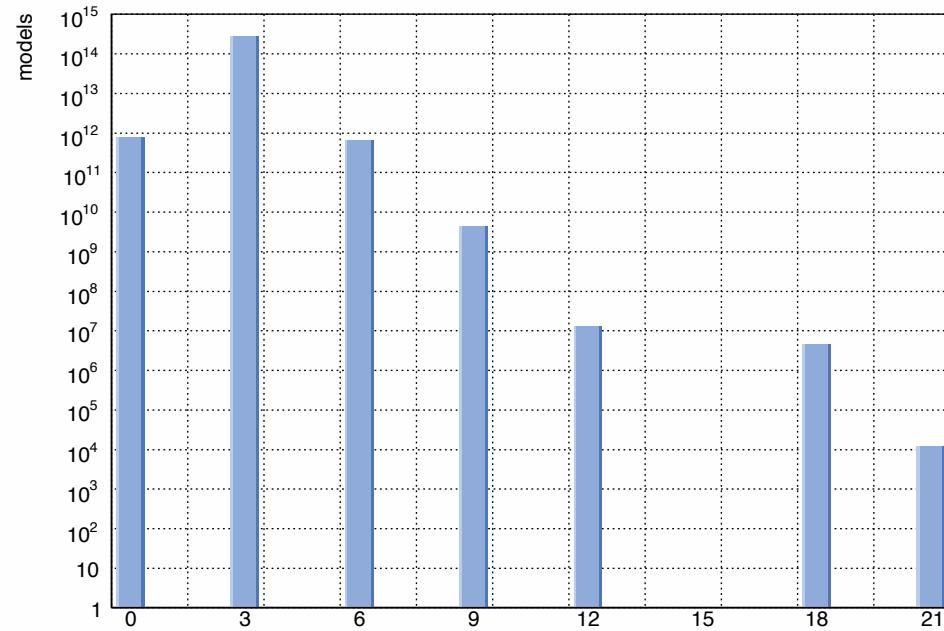
Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:

Z6'-orient

M

)39)



SM gauge group, 3 generations with # of Higgses



How many orientifold models exist which come close to the (spectrum of the) MSSM?

(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

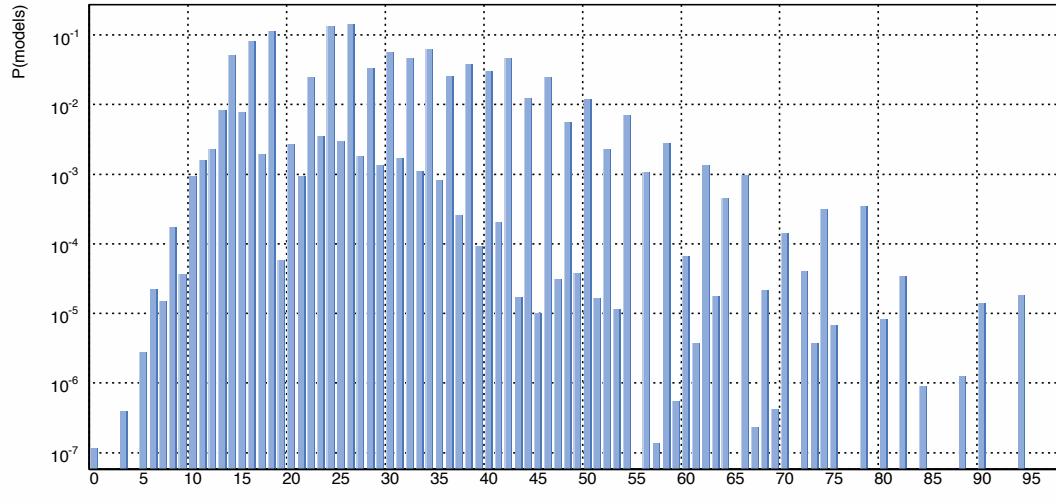
Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:

$Z_6'$  - . . . . .

39)



SM gauge group, 3 generations with # of chiral exotics

How many orientifold models exist which come close to the (spectrum of the) MSSM?

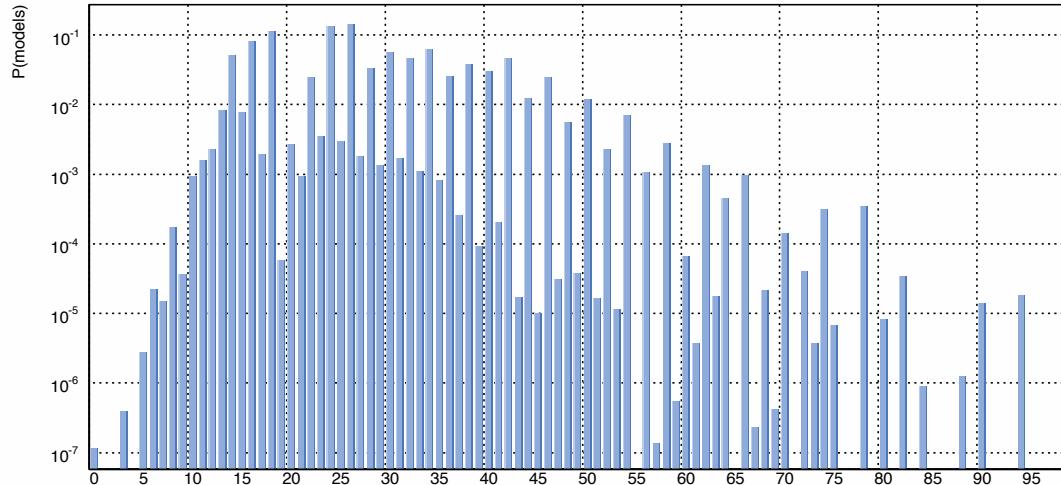
(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand; related work: Dijkstra, Huiszoon, Schellekens, hep-th/0411129; Anastasopoulos, Dijkstra, Kiritsis, Schellekens, hep-th/0605226; Douglas, Taylor, hep-th 0606109; Dienes, Lennek, hep-th/0610319)

Example:  $\mathcal{M}_6 = T^6 / (Z_N \times Z_M)$  IIA orientifold:

Systematic computer search (**NP complete problem**):

Look for solutions of a set of **diophantic equations**:

$Z_6'$  - . . . . .

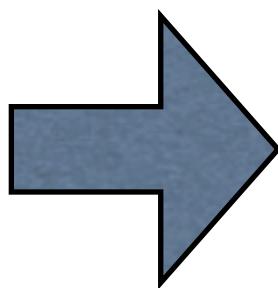


SM gauge group, 3 generations with # of chiral exotics

39)

ISB models with no chiral exotics are possible!

# Outline

- 
- Intersecting D-brane models
  - Mass scales in D-brane models
  - Stringy amplitudes for the LHC

(The LHC string hunter's companion)

### III) Mass scales in D-brane models:

### III) Mass scales in D-brane models:

There are 3 basic mass scales in D-brane compactifications:

### III) Mass scales in D-brane models:

There are 3 basic mass scales in D-brane compactifications:

**String scale:**

$$(1) : M_s = \frac{1}{\sqrt{\alpha'}}$$

### III) Mass scales in D-brane models:

There are 3 basic mass scales in D-brane compactifications:

**String scale:**

$$(1) : M_s = \frac{1}{\sqrt{\alpha'}}$$

**Compactification scale:** (2) :  $M_6 = \frac{1}{V_6^{1/6}}$

### III) Mass scales in D-brane models:

There are 3 basic mass scales in D-brane compactifications:

String scale:

$$(1) : M_s = \frac{1}{\sqrt{\alpha'}}$$

Compactification scale: (2) :  $M_6 = \frac{1}{V_6^{1/6}}$

Scale of wrapped D( $p+3$ )-branes (e.g. IIB:  $p=0,4$ ),  
(IIA:  $p=3$ ):

$$(3) : M_p^{\parallel} = \frac{1}{(V_p^{\parallel})^{1/p}}, \quad (3') : M_{6-p}^{\perp} = \frac{1}{(V_{6-p}^{\perp})^{1/(6-p)}}$$

$$V_6 = V_p^{\parallel} V_{6-p}^{\perp}$$

# There are 2 basic 4D observables:

There are 2 basic 4D observables:

Strength of 4D gravitational interactions:

$$(A) : M_{\text{Planck}}^2 \simeq M_s^8 V_6 \simeq 10^{19} \text{ GeV}$$

There are 2 basic 4D observables:

Strength of 4D gravitational interactions:

$$(A) : \quad M_{\text{Planck}}^2 \simeq M_s^8 V_6 \simeq 10^{19} \text{ GeV}$$

Strength of 4D gauge interactions:

$$(B) : \quad g_{Dp}^{-2} \simeq M_s^p V_p^{\parallel} \simeq \mathcal{O}(1)$$

$$\implies (V_p^{\parallel})^{-1/p} \simeq M_s$$

There are 2 basic 4D observables:

Strength of 4D gravitational interactions:

$$(A) : \quad M_{\text{Planck}}^2 \simeq M_s^8 V_6 \simeq 10^{19} \text{ GeV}$$

Strength of 4D gauge interactions:

$$(B) : \quad g_{Dp}^{-2} \simeq M_s^p V_p^{\parallel} \simeq \mathcal{O}(1)$$

$$\implies (V_p^{\parallel})^{-1/p} \simeq M_s$$

(A) and (B): leave one free parameter.

$M_s$  is a free parameter in D-brane compactifications !

There are 4 natural scenarios for the string scale:

There are 4 natural scenarios for the string scale:

(o) Planck scale scenario:

$M_s$  is the gravitational 4D Planck scale

$$M_s \equiv M_{\text{Planck}} \simeq 10^{19} \text{ GeV}$$

Gauge coupling unification at the Planck scales needs further effects (string threshold corrections, ...)

Alternatively relate the string scale to particles physics mass scales.

(i) GUT scale scenario:

$M_s$  is the 4D scale of gauge coupling unification

$$M_s \equiv M_{GUT} \simeq 10^{16} \text{ GeV}$$

$$M_{GUT} = M_{SM} \exp \left( \frac{g_{Dp}^{-2}(M_{SM}) - g_{Dp}^{-2}(M_{GUT})}{b_p} \right)$$

Recent GUT string model building in F-theory and IIB orientifolds:  
(Beasley, Heckman, Marsano, Saulina, Schafer-Nameki, Vafa; Donagi, Wijnholt; Blumenhagen, Braun, Grimm, Weigand; Andreas, Curio)

- D7-branes wrapped on del Pezzo surfaces
- GUT gauge group is broken by  $U(1)_Y$  flux

## (ii) SUSY breaking scenario:

$M_s$  is the intermediate 4D scale of supersymmetry breaking

(Balasubramanian, Conlon, Quevedo, Suruliz, ...)

$$M_s \equiv M_{SUSY} \simeq 10^{11} \text{ GeV}$$

Gravity mediation:

$$M_{SUSY} \sim \sqrt{M_{SM} M_{\text{Planck}}}$$

(No natural gauge coupling unification!)

### (iii) Low string scale scenario:

(Antoniadis, Arkani-Hamed, Dimopoulos, Dvali)

$M_s$  is the Standard Model (TeV) scale:

$$M_s \equiv M_{SM} \simeq 10^3 \text{ GeV}$$

(No natural gauge coupling unification!)

### (iii) Low string scale scenario:

(Antoniadis, Arkani-Hamed, Dimopoulos, Dvali)

$M_s$  is the Standard Model (TeV) scale:

$$M_s \equiv M_{SM} \simeq 10^3 \text{ GeV}$$

(No natural gauge coupling unification!)

## SUMMARY:

Table 1: The three different mass scales in D-brane models

	$M_s$ (GeV)	$L_s$ (m)	$M_6 = V_6^{-1/6}$ (GeV)	$V_6^{1/6}$ (m)	$M_2^\perp = (V_2^\perp)^{-1/2}$ (GeV)	$(V_2^\perp)^{1/2}$ (m)
(o)	$10^{19}$	$10^{-35}$	$10^{19}$	$10^{-35}$	$10^{19}$	$10^{-35}$
(i)	$10^{16}$	$10^{-32}$	$10^{15}$	$10^{-31}$	$10^{13}$	$10^{-29}$
(ii)	$10^{11}$	$10^{-27}$	$10^{6-7}$	$10^{-(22-23)}$	$10^3$	$10^{-19}$
(iii)	$10^3$	$10^{-19}$	$10^{-14/6}$	$10^{-14}$	$10^{-13}$	$10^{-3}$

Dimensionless volume in string units:

$$V'_6 = V_6 M_s^6 = \frac{M_{\text{Planck}}^2}{M_s^2} = 1, 10^6, 10^{16}, 10^{32}$$

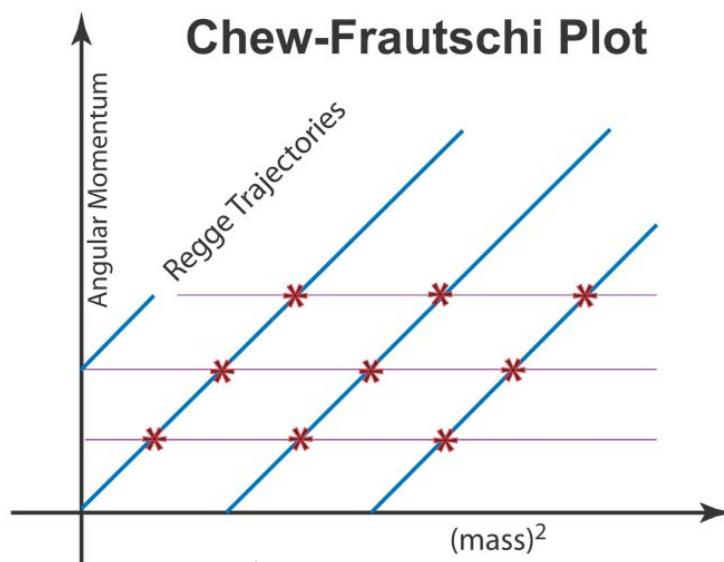
# There are 3 generic type of particles:

# There are 3 generic type of particles:

## (i) Stringy Regge excitations:

$$M_{\text{Regge}} = M_s = \frac{M_{\text{Planck}}}{\sqrt{V'_6}}$$

Open string excitations: completely universal (model independent), carry SM gauge quantum numbers



$$M_n^2 = M_s^2 \left( \sum_{k=1}^n \alpha_{-k}^\mu \alpha_k^\nu - 1 \right) = (n-1) M_s^2, \quad (n = 1, \dots, \infty)$$

## (ii) Overall volume modulus:

$$M_T = \frac{M_{\text{Planck}}}{(V'_6)^{3/2}} = 10^{19}, 10^{10}, 10^{-5}, 10^{-29} \text{ GeV}$$

Closed string, model independent, neutral under the SM,  
interacts only gravitationally

Problem: the very light mass causes a fifth force.

Would rule out TeV string scale !

## (ii) Overall volume modulus:

$$M_T = \frac{M_{\text{Planck}}}{(V'_6)^{3/2}} = 10^{19}, 10^{10}, 10^{-5}, 10^{-29} \text{ GeV}$$

Closed string, model independent, neutral under the SM,  
interacts only gravitationally

Problem: the very light mass causes a fifth force.

Would rule out TeV string scale !

## (ii) Overall volume modulus:

$$M_T = \frac{M_{\text{Planck}}}{(V'_6)^{3/2}} = 10^{19}, 10^{10}, 10^{-5}, 10^{-29} \text{ GeV}$$

Closed string, model independent, neutral under the SM,  
interacts only gravitationally

Problem: the very light mass causes a fifth force.

Would rule out TeV string scale !

But one expects a mass shift by radiative corrections:

$$\Delta M_T \simeq \frac{\langle T_\mu^\mu T_\mu^\mu \rangle}{M_{\text{Planck}}^2} \simeq \frac{M_s^4}{M_{\text{Planck}}^2} \simeq 10^{-13} \text{ GeV}$$

### (iii) D-brane cycle Kaluza Klein excitations:

$$M_{KK}^{\parallel} = \frac{1}{(V_p^{\parallel})^{1/p}} \simeq M_s = \frac{M_{\text{Planck}}}{(V'_6)^{1/2}}$$

Open strings, depend on the details of the internal geometry, carry SM gauge quantum numbers

### (iii) D-brane cycle Kaluza Klein excitations:

$$M_{KK}^{\parallel} = \frac{1}{(V_p^{\parallel})^{1/p}} \simeq M_s = \frac{M_{\text{Planck}}}{(V'_6)^{1/2}}$$

Open strings, depend on the details of the internal geometry, carry SM gauge quantum numbers

### SUMMARY:

The string Regge excitations (i) and the D-brane cycle KK modes (iii) are charged under the SM and have mass of order  $M_s$   $\Rightarrow$  can they be seen at LHC ?!

# Type IIB orientifolds: Realization of low string scale compatifications on „Swiss Cheese“ Manifolds:

(Abdussalam, Allanach, Balasubramanian, Berglund, Cicoli, Conlon, Kom, Quevedo, Suruliz;  
Blumenhagen, Moster, Plauschinn;

for model building and phenomenological aspects see: Conlon, Maharana, Quevedo, arXiv:0810.5660)

## Moduli potential:

Kähler potential:  $K = K_{cs} - 2 \log\left(V_6 + \frac{\xi}{2g_s^{\frac{3}{2}}}\right)$  (Becker, Becker,  
Haack, Louis)

Superpotential:  $W = W_{cs} + \sum A_i \exp(-a_i t_i)$

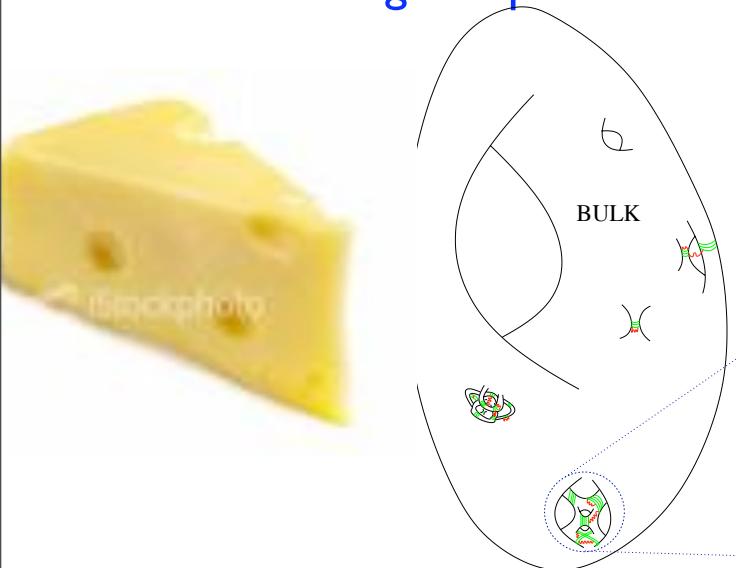
Moduli stabilization ➤

Minima: Large hierarchical scales with  $V_6 M_s^6 = 10^{16}, 10^{32}$

# Type IIB orientifolds: Realization of low string scale compatifications on „Swiss Cheese“ Manifolds:

(Abdussalam, Allanach, Balasubramanian, Berglund, Cicoli, Conlon, Kom, Quevedo, Suruliz;  
Blumenhagen, Moster, Plauschinn;

for model building and phenomenological aspects see: Conlon, Maharana, Quevedo, arXiv:0810.5660)



## 2 requirements:

- Negative Euler number.
- SM lives on D7-branes around small cycles of the CY. One needs at least one blow-up mode (resolves point like singularity).

## Moduli potential:

Kähler potential:  $K = K_{cs} - 2 \log \left( V_6 + \frac{\xi}{2g_s^{\frac{3}{2}}} \right)$  (Becker, Becker, Haack, Louis)

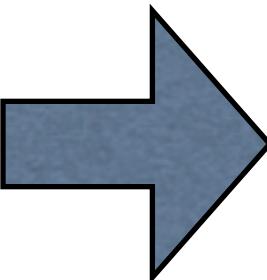
Superpotential:  $W = W_{cs} + \sum A_i \exp(-a_i t_i)$

## Moduli stabilization ➤

Minima: Large hierarchical scales with  $V_6 M_s^6 = 10^{16}, 10^{32}$

# Outline

- Intersecting D-brane models
- Mass scales in D-brane models
- Stringy amplitudes for the LHC



(The LHC string hunter's companion)

(D. Lüst, S. Stieberger, T. Taylor, arXiv:0807.3333; L. Anchordoqui, H. Goldberg, D. Lüst, S. Nawata, S. Stieberger, T. Taylor, arXiv:0808.0497 [hep-ph];  
L. Anchordoqui, H. Goldberg, D. Härtl, D. Lüst, S. Nawata, O. Schlotterer, S. Stieberger, T. Taylor, to appear)

# III) The LHC String Hunter's Companion:

### **III) The LHC String Hunter's Companion:**

**Test of D-brane models at the LHC:**

**(New stringy physics of beyond the SM)**

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

- Massive extra (anomalous)  $Z'$  U(1) gauge bosons

(also kinetic mixing of  $Z'$  with photon  
and milli-charged particles)

(Abel, Goodsell, Jäckel, Khoze,  
Ringwald, arXiv: 0803.1449)

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

- Massive extra (anomalous)  $Z'$  U(1) gauge bosons

(also kinetic mixing of  $Z'$  with photon  
and milli-charged particles)

(Abel, Goodsell, Jäckel, Khoze,  
Ringwald, arXiv: 0803.1449)

- Massive black holes (for strong string coupling)

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

- Massive extra (anomalous)  $Z'$  U(1) gauge bosons

(also kinetic mixing of  $Z'$  with photon  
and milli-charged particles)

(Abel, Goodsell, Jäckel, Khoze,  
Ringwald, arXiv: 0803.1449)

- Massive black holes (for strong string coupling)
- Regge excitations of higher spin

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

- Massive extra (anomalous)  $Z'$  U(1) gauge bosons

(also kinetic mixing of  $Z'$  with photon  
and milli-charged particles)

(Abel, Goodsell, Jäckel, Khoze,  
Ringwald, arXiv: 0803.1449)

- Massive black holes (for strong string coupling)
- Regge excitations of higher spin
- Kaluza Klein (KK) (and winding) modes

### III) The LHC String Hunter's Companion:

Test of D-brane models at the LHC:

(New stringy physics of beyond the SM)

New massive particles at string scale  $M_s$  :

- Massive extra (anomalous)  $Z'$  U(1) gauge bosons

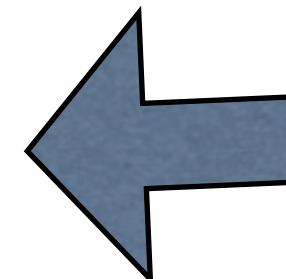
(also kinetic mixing of  $Z'$  with photon  
and milli-charged particles)

(Abel, Goodsell, Jäckel, Khoze,  
Ringwald, arXiv: 0803.1449)

- Massive black holes (for strong string coupling)

- Regge excitations of higher spin

- Kaluza Klein (KK) (and winding) modes

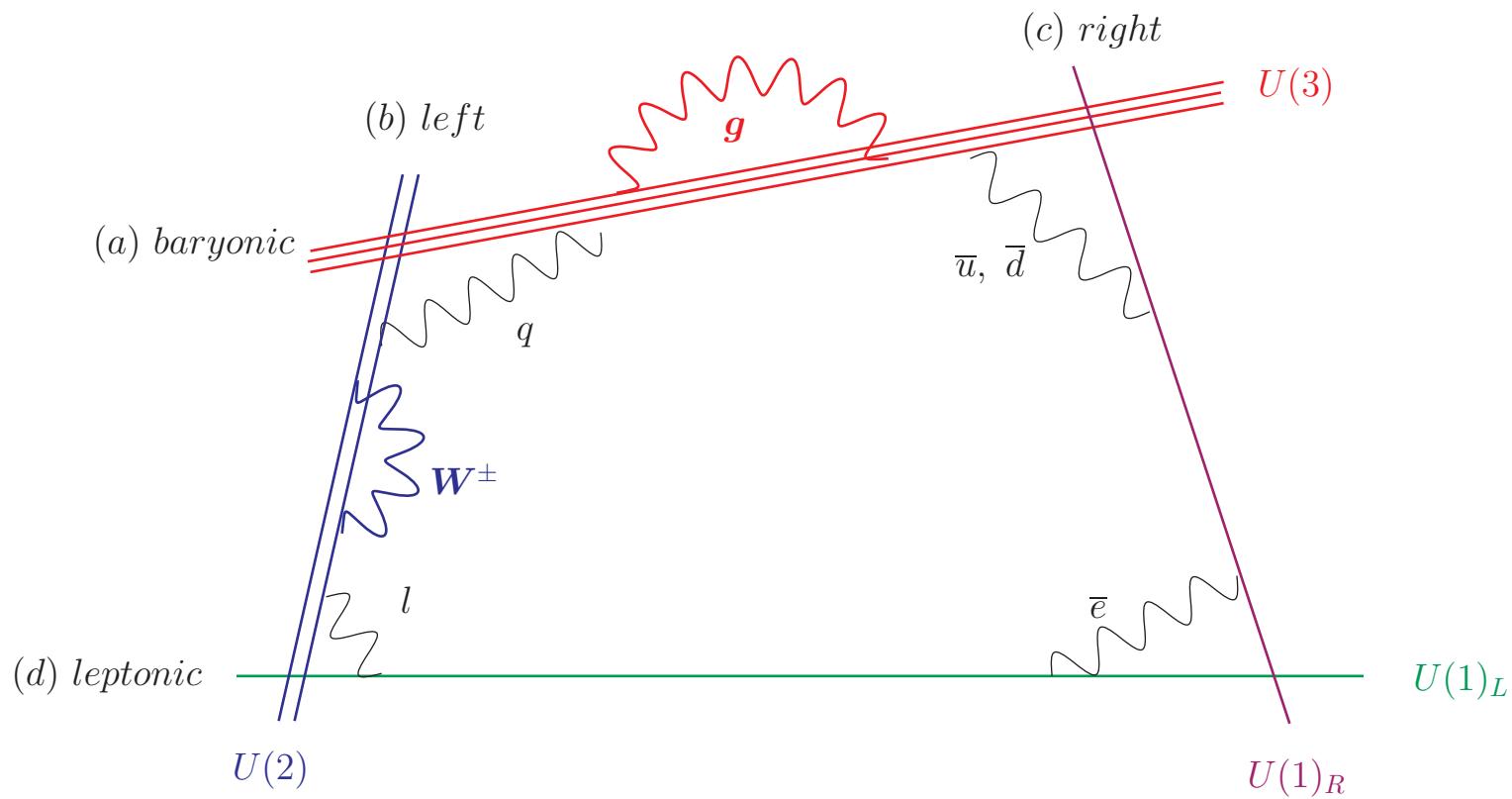


One has to compute the parton model cross sections of  
SM fields into new stringy states !

# Parton model cross sections of SM-fields:

Disk amplitude among n external SM fields ( $q, l, g, \gamma, Z^0, W^\pm$ ):

$$n = 4 : \mathcal{A}(\Phi^1, \Phi^2, \Phi^3, \Phi^4) = < V_{\Phi^1}(z_1) V_{\Phi^2}(z_2) V_{\Phi^3}(z_3) V_{\Phi^4}(z_4) >_{disk}$$



**Parton model cross sections of SM-fields:**

**Disk amplitude among n external SM fields** ( $q, l, g, \gamma, Z^0, W^\pm$ ):

$$n = 4 : \quad \mathcal{A}(\Phi^1, \Phi^2, \Phi^3, \Phi^4) = < V_{\Phi^1}(z_1) \ V_{\Phi^2}(z_2) \ V_{\Phi^3}(z_3) \ V_{\Phi^4}(z_4) >_{disk}$$

**These amplitudes are dominated by the following poles:**

Parton model cross sections of SM-fields:

Disk amplitude among n external SM fields  $(q, l, g, \gamma, Z^0, W^\pm)$ :

$$n = 4 : \quad \mathcal{A}(\Phi^1, \Phi^2, \Phi^3, \Phi^4) = < V_{\Phi^1}(z_1) \ V_{\Phi^2}(z_2) \ V_{\Phi^3}(z_3) \ V_{\Phi^4}(z_4) >_{disk}$$

These amplitudes are dominated by the following poles:

- Exchange of SM fields

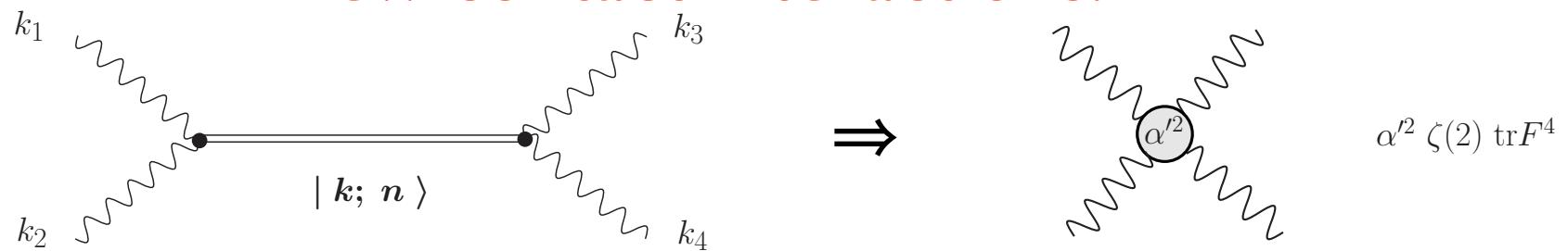
# Parton model cross sections of SM-fields:

Disk amplitude among n external SM fields ( $q, l, g, \gamma, Z^0, W^\pm$ ):

$$n = 4 : \mathcal{A}(\Phi^1, \Phi^2, \Phi^3, \Phi^4) = \langle V_{\Phi^1}(z_1) V_{\Phi^2}(z_2) V_{\Phi^3}(z_3) V_{\Phi^4}(z_4) \rangle_{disk}$$

These amplitudes are dominated by the following poles:

- Exchange of SM fields
- Exchange of string Regge resonances (Veneziano like ampl.)  
⇒ new contact interactions:



$$\mathcal{A}(k_1, k_2, k_3, k_4; \alpha') \sim -\frac{\Gamma(-\alpha' s) \Gamma(1 - \alpha' u)}{\Gamma(-\alpha' s - \alpha' u)} = \sum_{n=0}^{\infty} \frac{\gamma(n)}{s - M_n^2} \sim \frac{t}{s} - \frac{\pi^2}{6} tu (\alpha')^2 + \dots$$

$$V_t(\alpha') = \frac{\Gamma(1 - s/M_{\text{string}}^2) \Gamma(1 - u/M_{\text{string}}^2)}{\Gamma(1 - t/M_{\text{string}}^2)} = 1 - \frac{\pi^2}{6} M_{\text{string}}^{-4} su - \zeta(3) M_{\text{string}}^{-6} stu + \dots \rightarrow 1|_{\alpha' \rightarrow 0}$$

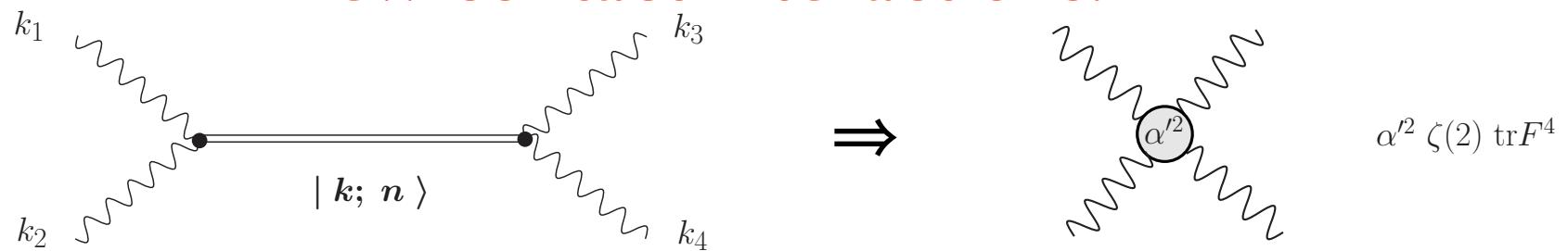
# Parton model cross sections of SM-fields:

Disk amplitude among n external SM fields ( $q, l, g, \gamma, Z^0, W^\pm$ ):

$$n = 4 : \mathcal{A}(\Phi^1, \Phi^2, \Phi^3, \Phi^4) = \langle V_{\Phi^1}(z_1) V_{\Phi^2}(z_2) V_{\Phi^3}(z_3) V_{\Phi^4}(z_4) \rangle_{disk}$$

These amplitudes are dominated by the following poles:

- Exchange of SM fields
- Exchange of string Regge resonances (Veneziano like ampl.)  
⇒ new contact interactions:



$$\mathcal{A}(k_1, k_2, k_3, k_4; \alpha') \sim -\frac{\Gamma(-\alpha' s) \Gamma(1 - \alpha' u)}{\Gamma(-\alpha' s - \alpha' u)} = \sum_{n=0}^{\infty} \frac{\gamma(n)}{s - M_n^2} \sim \frac{t}{s} - \frac{\pi^2}{6} tu (\alpha')^2 + \dots$$

$$V_t(\alpha') = \frac{\Gamma(1 - s/M_{\text{string}}^2) \Gamma(1 - u/M_{\text{string}}^2)}{\Gamma(1 - t/M_{\text{string}}^2)} = 1 - \frac{\pi^2}{6} M_{\text{string}}^{-4} su - \zeta(3) M_{\text{string}}^{-6} stu + \dots \rightarrow 1|_{\alpha' \rightarrow 0}$$

- Exchange of KK and winding modes (model dependent)

The string scattering amplitudes exhibit some interesting properties:

- Interesting mathematical structure
- They go beyond the N=4 Yang-Mills amplitudes:

(i) The contain quarks & leptons in fundamental repr.

Quark, lepton vertex operators:

$$V_{q,l}(z, u, k) = u^\alpha S_\alpha(z) \Xi^{a \cap b}(z) e^{-\phi(z)/2} e^{ik \cdot X(z)}$$

Fermions: boundary changing (twist) operators!

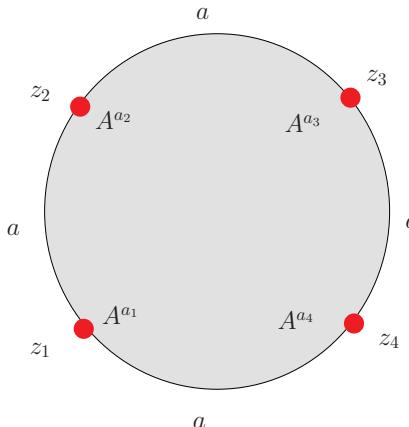
(ii) They contain stringy corrections.

# Striking relation between quark and gluon amplitudes:

## (i) Four point scattering amplitudes (2 jet events):

4 gluons:

(Stieberger, Taylor)



Field theory factors:

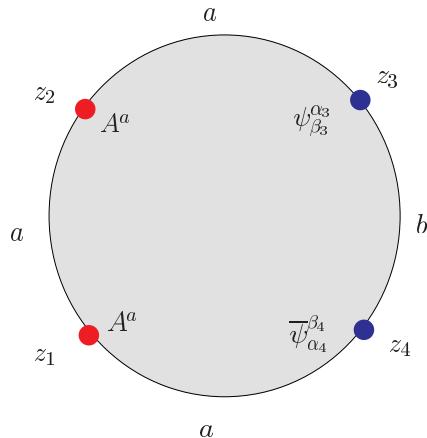
$$\mathcal{M}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

$$\langle ij \rangle = (\lambda_i)^\alpha (\lambda_j)_\alpha$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+) = V^{(4)}(\alpha', k_i) \times \mathcal{M}_{\text{YM}}^{(4)}$$

2 gluons, 2 quarks:

(Lüst, Stieberger, Taylor)



$$\mathcal{N}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 14 \rangle \langle 13 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

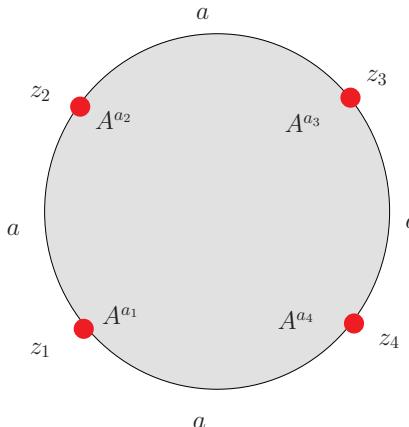
$$\mathcal{A}(g_1^-, g_2^+, q_3^-, \bar{q}_4^+) = V^{(4)}(\alpha', k_i) \times \mathcal{N}_{\text{YM}}^{(4)}$$

# Striking relation between quark and gluon amplitudes:

## (i) Four point scattering amplitudes (2 jet events):

4 gluons:

(Stieberger, Taylor)



Field theory factors:

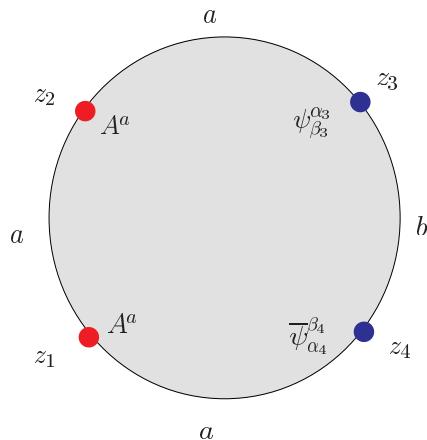
$$\mathcal{M}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

$$\langle ij \rangle = (\lambda_i)^\alpha (\lambda_j)_\alpha$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{M}_{\text{YM}}^{(4)}, \quad (V^{(4)} = 1 + \zeta(2)\mathcal{O}(\alpha'^2))$$

2 gluons, 2 quarks:

(Lüst, Stieberger, Taylor)



$$\mathcal{N}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 14 \rangle \langle 13 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

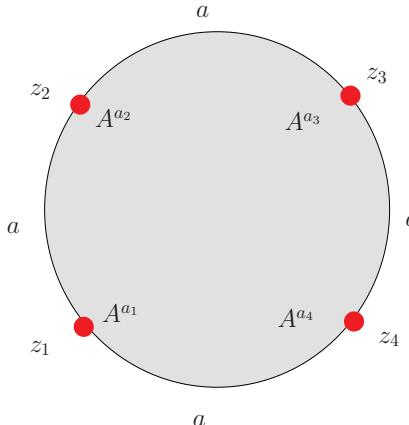
$$\mathcal{A}(g_1^-, g_2^+, q_3^-, \bar{q}_4^+) = V^{(4)}(\alpha', k_i) \times \mathcal{N}_{\text{YM}}^{(4)}$$

# Striking relation between quark and gluon amplitudes:

## (i) Four point scattering amplitudes (2 jet events):

4 gluons:

(Stieberger, Taylor)



Field theory factors:

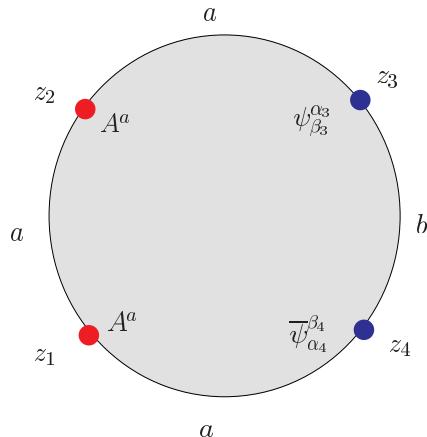
$$\mathcal{M}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

$$\langle ij \rangle = (\lambda_i)^\alpha (\lambda_j)_\alpha$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{M}_{\text{YM}}^{(4)}, \quad (V^{(4)} = 1 + \zeta(2)\mathcal{O}(\alpha'^2))$$

2 gluons, 2 quarks:

(Lüst, Stieberger, Taylor)



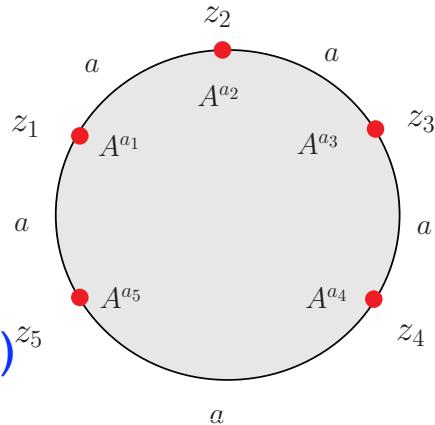
$$\mathcal{N}_{\text{YM}}^{(4)} = \frac{4g_{\text{YM}}^2 \langle 14 \rangle \langle 13 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^+, q_3^-, \bar{q}_4^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{N}_{\text{YM}}^{(4)}$$

## (ii) Five point scattering amplitudes (3 jet events):

5 gluons:

(Stieberger, Taylor (2006))



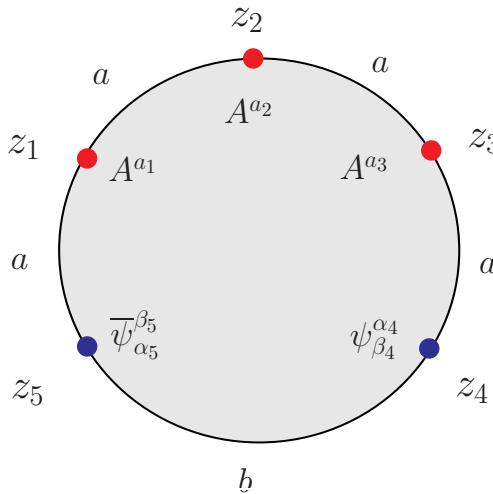
Field theory factors:

$$\mathcal{M}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+, g_5^+) = (V^{(5)}(\alpha', k_i) - 2i\epsilon(1, 2, 3, 4)P^{(5)}(\alpha', k_i)) \times \mathcal{M}_{\text{YM}}^{(5)}$$

3 gluons, 2 quarks:

(D. Lüst, O. Schlotterer,  
S. Stieberger, T. Taylor, work in  
progress).



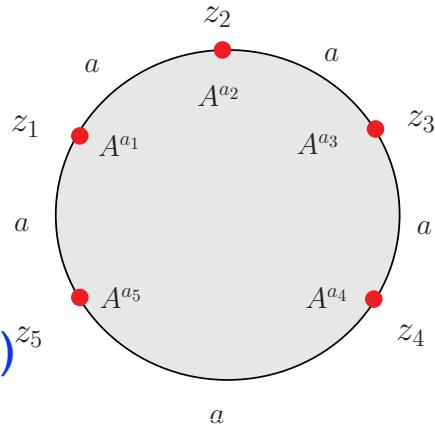
$$\mathcal{N}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 15 \rangle \langle 14 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^+, g_3^+, q_4^-, \bar{q}_5^+) = (V^{(5)}(\alpha', k_i) - 2i\epsilon(1, 2, 3, 4)P^{(5)}(\alpha', k_i)) \times \mathcal{N}_{\text{YM}}^{(5)}$$

## (ii) Five point scattering amplitudes (3 jet events):

5 gluons:

(Stieberger, Taylor (2006))



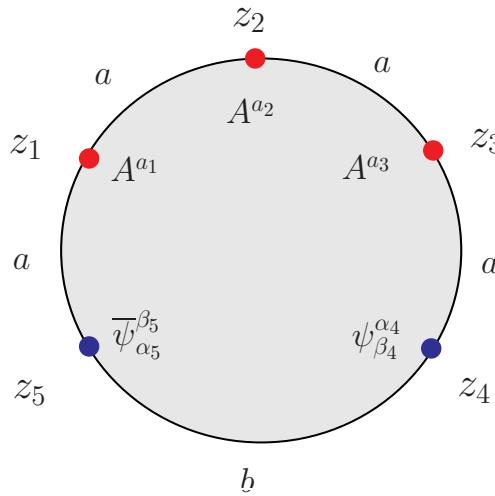
Field theory factors:

$$\mathcal{M}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+, g_5^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{M}_{\text{YM}}^{(5)}, \quad (V^{(5)} = 1 + \zeta(2)\mathcal{O}(\alpha'^2), \quad P^{(5)} = \zeta(2)\mathcal{O}(\alpha'^2))$$

3 gluons, 2 quarks:

(D. Lüst, O. Schlotterer,  
S. Stieberger, T. Taylor, work in  
progress).



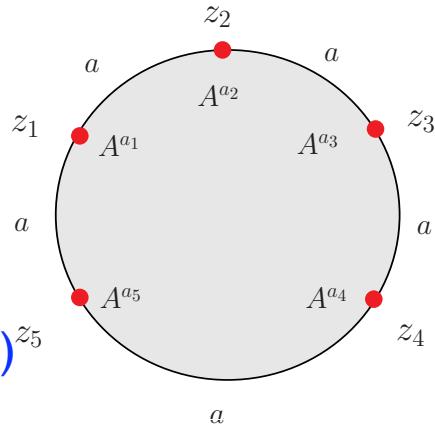
$$\mathcal{N}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 15 \rangle \langle 14 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^+, g_3^+, q_4^-, \bar{q}_5^+) = (V^{(5)}(\alpha', k_i) - 2i\epsilon(1, 2, 3, 4)P^{(5)}(\alpha', k_i)) \times \mathcal{N}_{\text{YM}}^{(5)}$$

## (ii) Five point scattering amplitudes (3 jet events):

5 gluons:

(Stieberger, Taylor (2006))



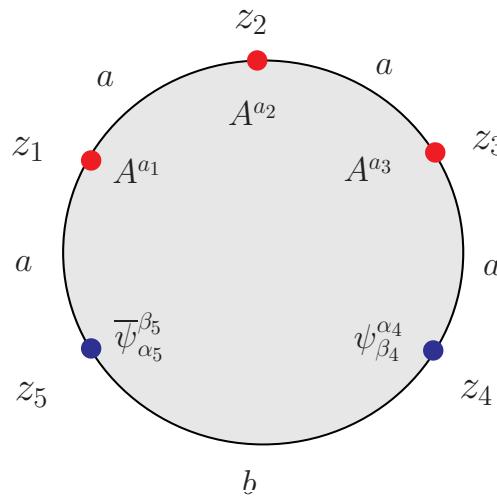
Field theory factors:

$$\mathcal{M}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^-, g_3^+, g_4^+, g_5^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{M}_{\text{YM}}^{(5)}, \quad (V^{(5)} = 1 + \zeta(2)\mathcal{O}(\alpha'^2), \quad P^{(5)} = \zeta(2)\mathcal{O}(\alpha'^2))$$

3 gluons, 2 quarks:

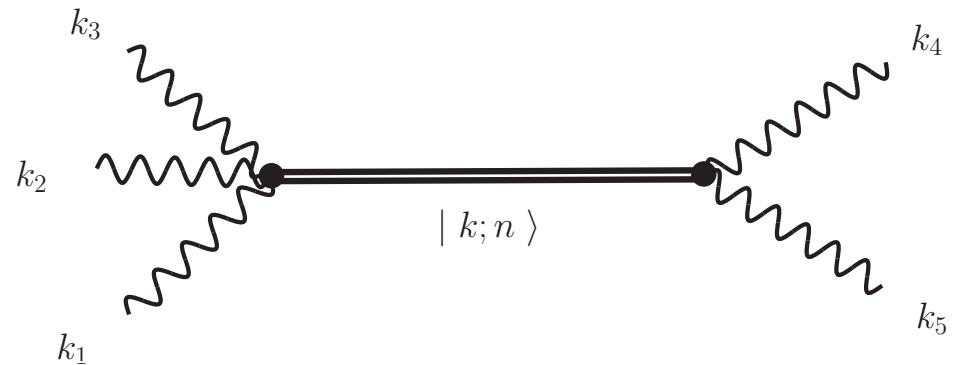
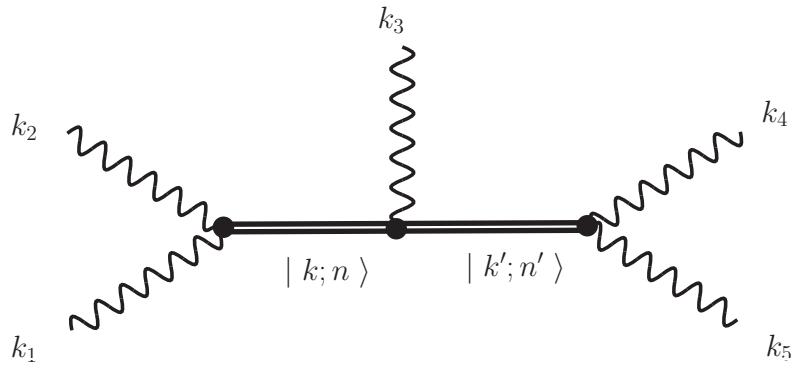
(D. Lüst, O. Schlotterer,  
S. Stieberger, T. Taylor, work in  
progress).



$$\mathcal{N}_{\text{YM}}^{(5)} = \frac{4g_{\text{YM}}^3 \langle 15 \rangle \langle 14 \rangle^3}{\langle 12 \rangle \langle 23 \rangle \dots \langle 51 \rangle}$$

$$\mathcal{A}(g_1^-, g_2^+, g_3^+, q_4^-, \bar{q}_5^+)_{\alpha' \rightarrow 0} \rightarrow \mathcal{N}_{\text{YM}}^{(5)}$$

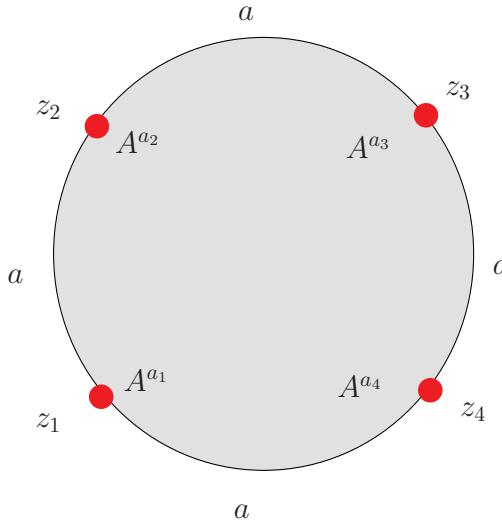
The two kinds of amplitudes are universal: the same Regge states are exchanged:



- n-point tree amplitudes with 0 or 2 open string fermions (quarks, leptons) and n or n-2 gauge bosons (gluons) are completely **model independent**.  
⇒ Information about the string Regge spectrum.

# 4 gauge boson amplitudes:

Disk amplitude:



Only string Regge resonances are exchanged  $\Rightarrow$

This amplitude is completely model independent!

Examples:

$$|\mathcal{A}(gg \rightarrow gg)|^2 = g_3^4 \left( \frac{1}{s^2} + \frac{1}{t^2} + \frac{1}{u^2} \right) \left[ \frac{9}{4} s^2 V_s^2(\alpha') - \frac{1}{3} (s V_s(\alpha'))^2 + (s \leftrightarrow t) + (s \leftrightarrow u) \right]$$

$\Rightarrow$  dijet events

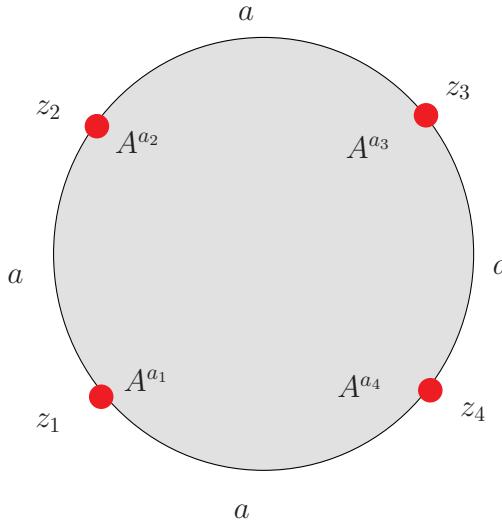
$$|\mathcal{A}(gg \rightarrow g\gamma(Z^0))|^2 = g_3^4 \frac{5}{6} Q_A^2 \left( \frac{1}{s^2} + \frac{1}{t^2} + \frac{1}{u^2} \right) (s V_s(\alpha') + t V_t(\alpha') + u V_u(\alpha'))^2$$

(Anchordoqui, Goldberg,  
Nawata, Taylor,  
arXiv:0712.0386)

Observable at LHC for  $M_{\text{string}} = 3$  TeV

# 4 gauge boson amplitudes:

Disk amplitude:



Only string Regge resonances are exchanged  $\Rightarrow$

This amplitude is completely model independent!

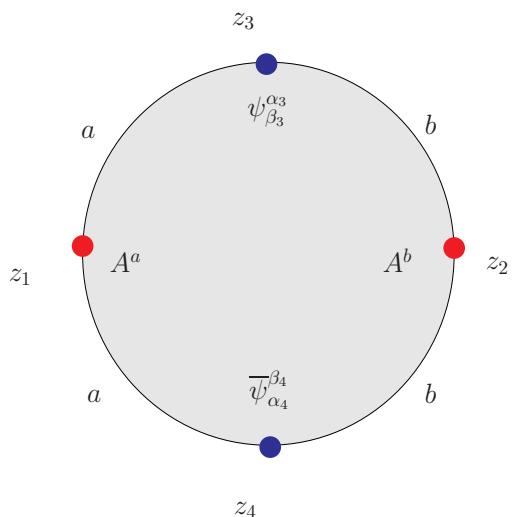
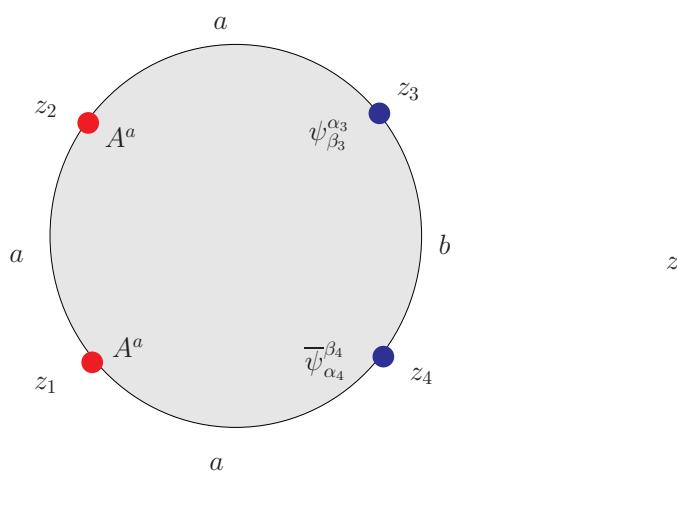
Examples:

$\alpha' \rightarrow 0$  : agreement with SM!

$$|\mathcal{A}(gg \rightarrow gg)|_{\alpha' \rightarrow 0}^2 \rightarrow \left( \frac{1}{s^2} + \frac{1}{t^2} + \frac{1}{u^2} \right) \frac{9}{4} (s^2 + t^2 + u^2)$$

$$|\mathcal{A}(gg \rightarrow \gamma(Z^0))|_{\alpha' \rightarrow 0}^2 \rightarrow 0$$

# 2 gauge boson - two fermion amplitude:



Note: Cullen, Perelstein, Peskin (2000)  
considered:

$$e^+ e^- \rightarrow \gamma\gamma$$

Only string Regge resonances are exchanged  $\Rightarrow$

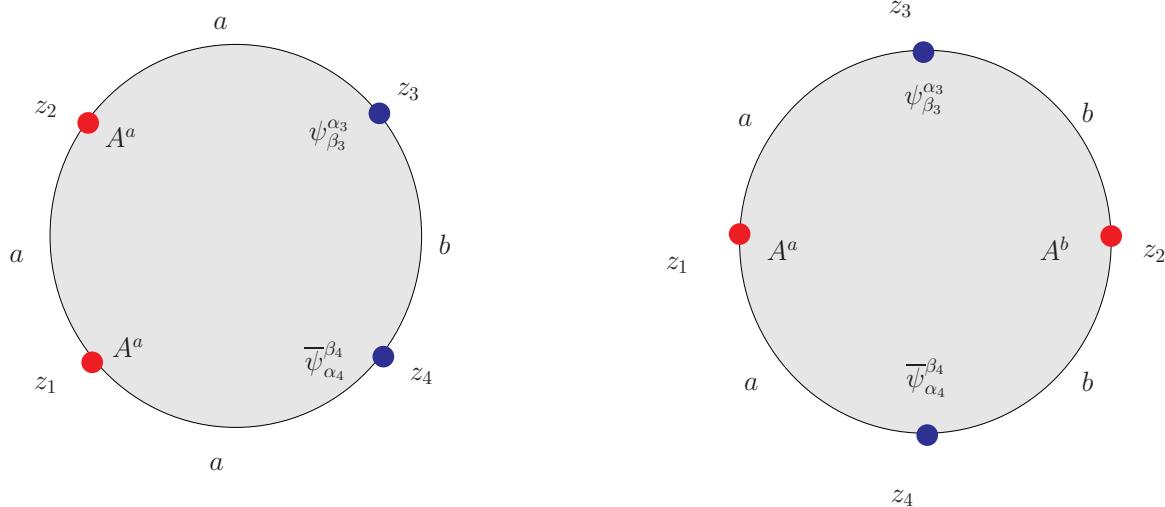
These amplitudes are completely model independent!

$$|\mathcal{A}(qg \rightarrow qg)|^2 = g_3^4 \frac{s^2 + u^2}{t^2} \left[ V_s(\alpha') V_u(\alpha') - \frac{4}{9} \frac{1}{su} (sV_s(\alpha') + uV_u(\alpha'))^2 \right]$$

$\Rightarrow$  dijet events

$$|\mathcal{A}(qg \rightarrow q\gamma(Z^0))|^2 = -\frac{1}{3} g_3^4 Q_A^2 \frac{s^2 + u^2}{sut^2} (sV_s(\alpha') + uV_u(\alpha'))^2$$

## 2 gauge boson - two fermion amplitude:



Note: Cullen, Perelstein, Peskin (2000)  
considered:

$$e^+ e^- \rightarrow \gamma\gamma$$

Only string Regge resonances are exchanged  $\Rightarrow$

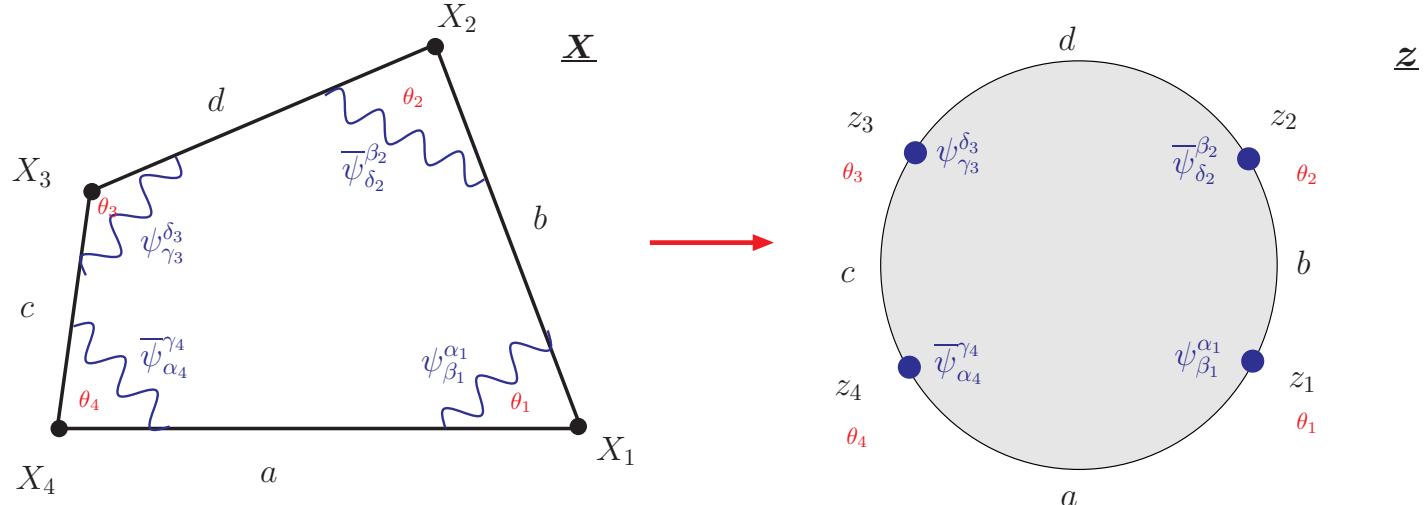
These amplitudes are completely model independent!

$\alpha' \rightarrow 0$  : agreement with SM !

$$|\mathcal{A}(qg \rightarrow qg)|_{\alpha' \rightarrow 0}^2 = g_3^4 \frac{s^2 + u^2}{t^2} \left[ 1 - \frac{4}{9} \frac{1}{su} (s + u)^2 \right]$$

$$|\mathcal{A}(qg \rightarrow q\gamma(Z^0))|_{\alpha' \rightarrow 0}^2 = -\frac{1}{3} g_3^4 Q_A^2 \frac{s^2 + u^2}{s u t^2} (s + u)^2$$

# 4 fermion amplitudes:



# Exchange of **Regge**, KK and winding resonances.

These amplitudes are more model dependent  
and test the internal CY geometry.

Constrained by FCNC's and/or proton decay.

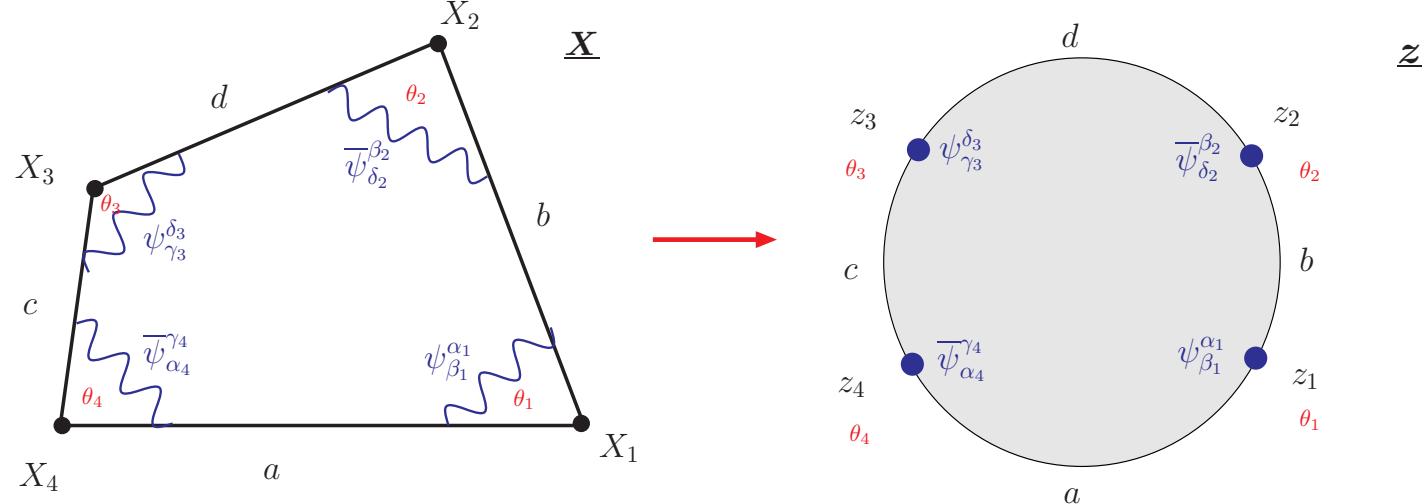
(Klebanov, Witten, hep-th/0304079; Abel, Lebedev, Santiago, hep-th/0312157)

E.g.

$$|\mathcal{A}(qq \rightarrow qq)|^2 = \frac{2}{9} \frac{1}{t^2} \left[ (sF_{tu}^{bb}(\alpha'))^2 + (sF_{tu}^{cc}(\alpha'))^2 + (uG_{ts}^{bc}(\alpha'))^2 + (uG_{ts}^{cb}(\alpha'))^2 \right] + \frac{2}{9} \frac{1}{u^2} \left[ (sF_{ut}^{bb}(\alpha'))^2 + (sF_{ut}^{cc}(\alpha'))^2 + (tG_{us}^{bc}(\alpha'))^2 + (tG_{us}^{cb}(\alpha'))^2 \right] - \frac{4}{27} \frac{s^2}{tu} F_{tu}^{bb}(\alpha') F_{ut}^{bb}(\alpha') + F_{tu}^{cc}(\alpha') F_{ut}^{cc}(\alpha')$$

depend on internal geometry

## 4 fermion amplitudes:



Exchange of **Regge, KK and winding resonances.**

These amplitudes are more model dependent  
and test the internal CY geometry.

Constrained by FCNC's and/or proton decay.

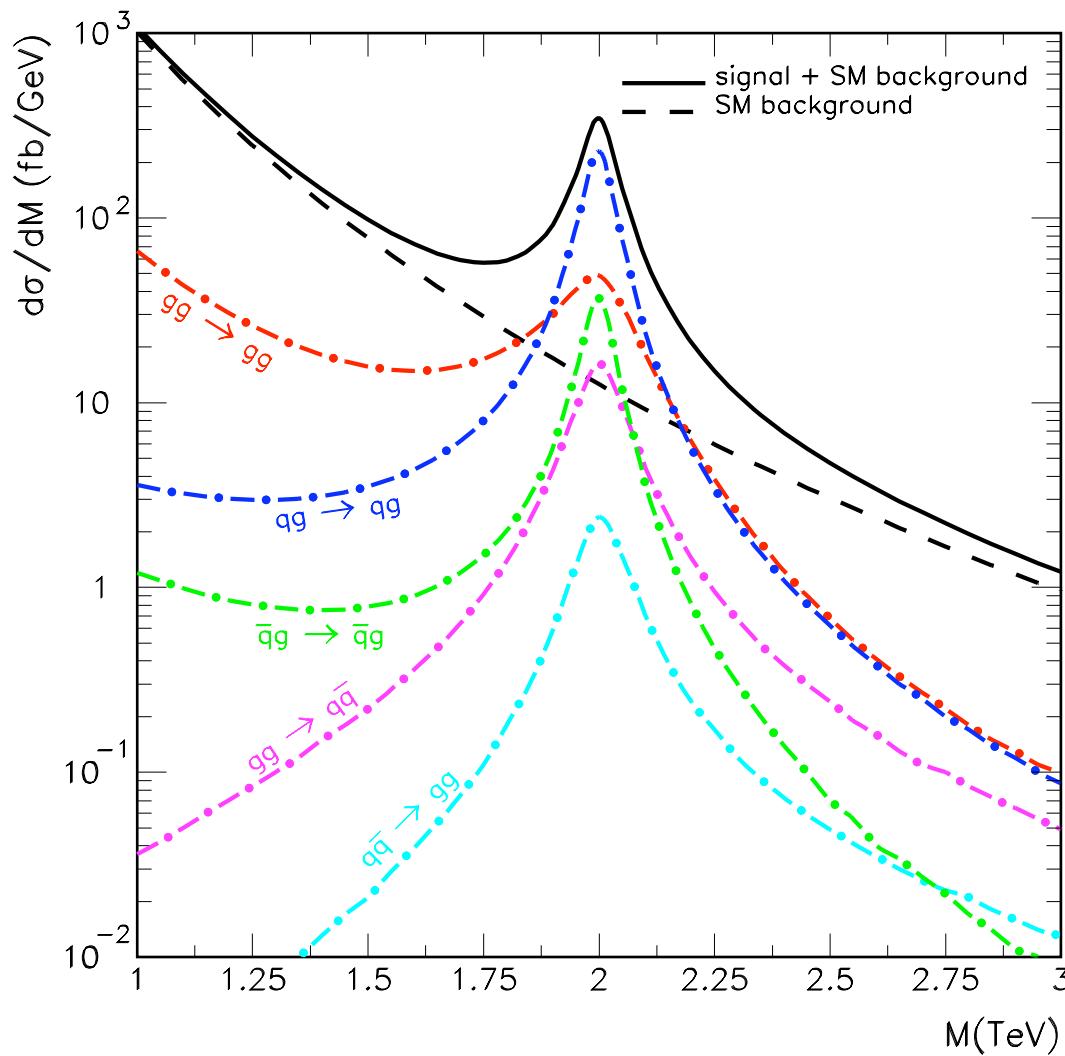
(Klebanov, Witten, hep-th/0304079; Abel, Lebedev, Santiago, hep-th/0312157)

E.g.

$\alpha' \rightarrow 0$  : agreement with SM !

$$|\mathcal{A}(qq \rightarrow qq)|_{\alpha' \rightarrow 0}^2 \rightarrow \frac{4}{9} \left[ \frac{s^2 + u^2}{t^2} \right] + \frac{4}{9} \left[ \frac{s^2 + t^2}{u^2} \right] - \frac{8}{27} \frac{s^2}{tu}$$

# These stringy corrections can be seen in dijet events at LHC:



(Anchordoqui, Goldberg, Lüst, Nawata,  
Stieberger, Taylor, arXiv:0808.0497[hep-ph])

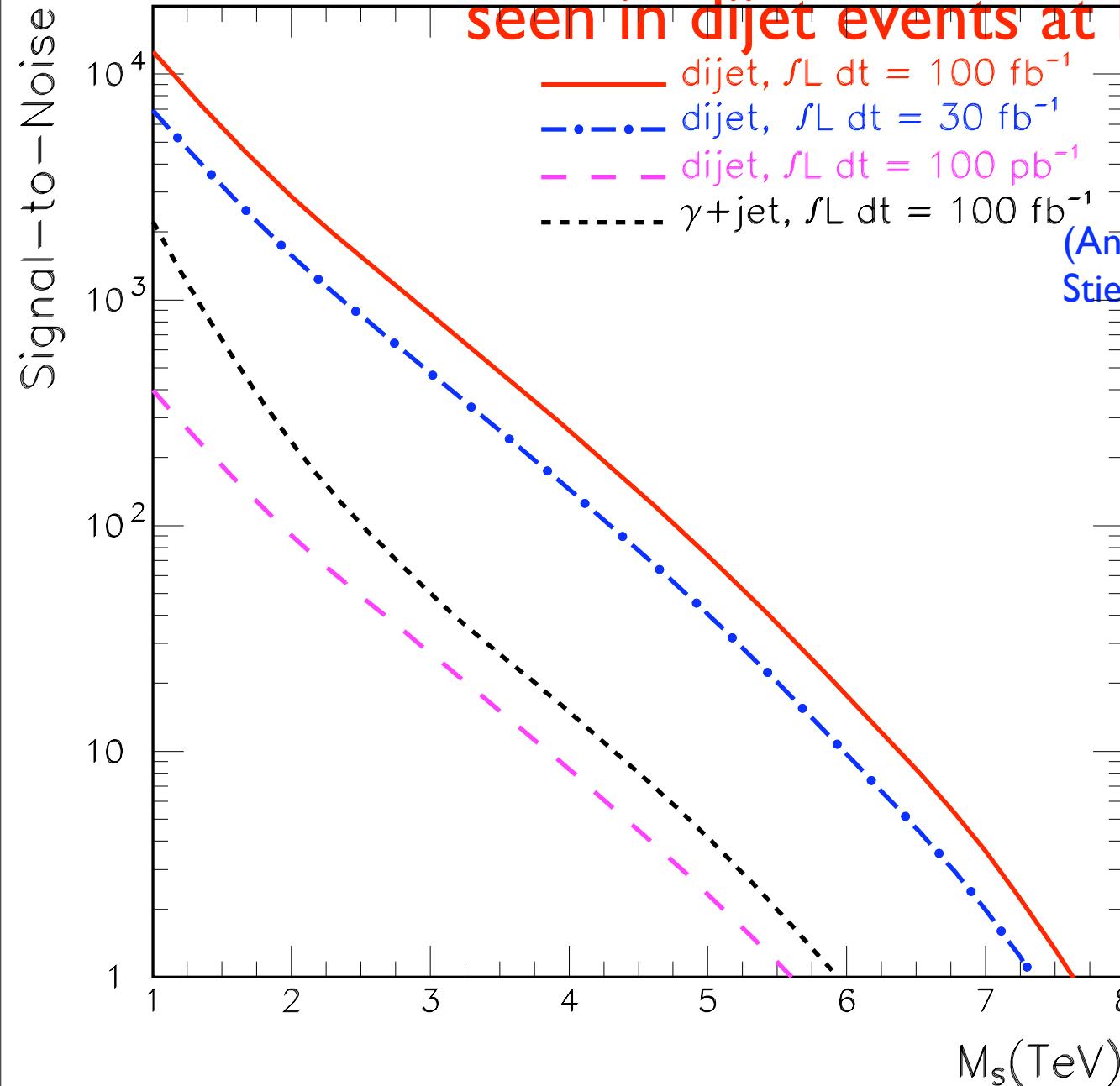
$$M_{\text{Regge}} = 2 \text{ TeV}$$

$$\Gamma_{\text{Regge}} = 15 - 150 \text{ GeV}$$

Widths can be computed in a  
model independent way !

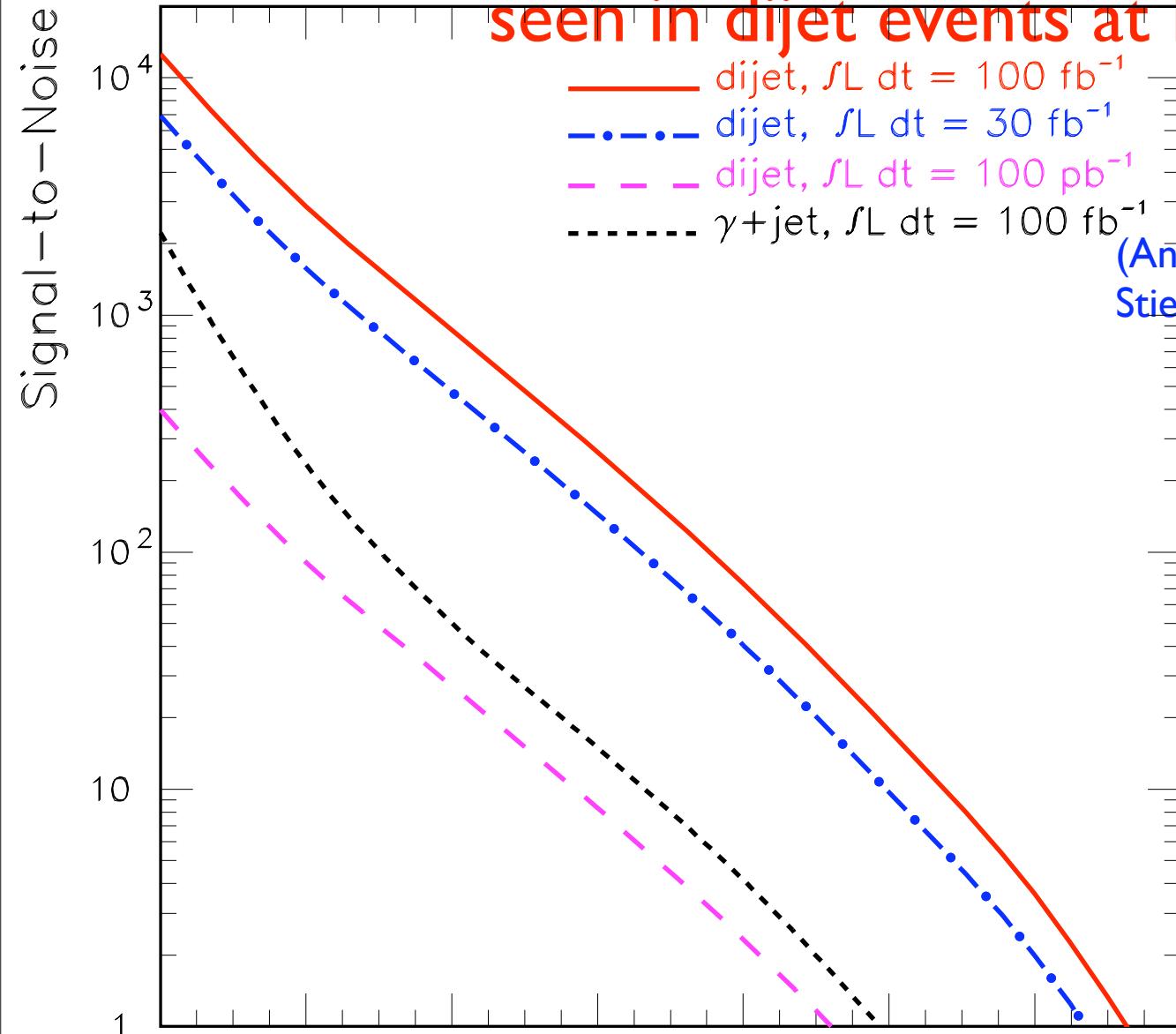
(Anchordoqui, Goldberg, Taylor,  
arXiv:0806.3420)

These stringy corrections can be  
seen in dijet events at LHC:



(Anchordoqui, Goldberg, Lüst, Nawata,  
Stieberger, Taylor, arXiv:0808.0497[hep-ph])

These stringy corrections can be  
seen in dijet events at LHC:



(Anchordoqui, Goldberg, Lüst, Nawata,  
Stieberger, Taylor, arXiv:0808.0497[hep-ph])

There are possible also stringy Drell-Yan processes like  
 $q\bar{q} \rightarrow l\bar{l}$

- KK modes are seen in scattering processes with more than 2 fermions.

⇒ Information about the internal geometry.

KK modes are exchanged in t- and u-channel processes and exhibit an interesting angular distribution.

(L.Anchordoqui, H. Goldberg, D. Lüst, S. Nawata, S. Stieberger, T. Taylor, paper in preparation)

# Conclusions

# Conclusions

- There exists many ISB models with SM like spectra without chiral exotics.

# Conclusions

- There exists many ISB models with SM like spectra without chiral exotics.
- One can make some model independent predictions:  
(Independent of amount of (unbroken) supersymmetry!)

String tree level, 4-point processes with 2 or 4 gluons

👉 observable at LHC ?? -  $M_{\text{string}}$  ??

# Conclusions

- There exists many ISB models with SM like spectra without chiral exotics.
- One can make some model independent predictions:  
**(Independent of amount of (unbroken) supersymmetry!)**  
String tree level, 4-point processes with 2 or 4 gluons  
👉 observable at LHC ?? -  $M_{\text{string}}$  ??

Computations done at weak string coupling !

Black holes are heavier than Regge states:  $M_{b.h.} = \frac{M_{\text{string}}}{g_{\text{string}}}$

# Conclusions

- There exists many ISB models with SM like spectra without chiral exotics.
- One can make some model independent predictions:  
(Independent of amount of (unbroken) supersymmetry!)

String tree level, 4-point processes with 2 or 4 gluons

👉 observable at LHC ?? -  $M_{\text{string}}$  ??

Computations done at weak string coupling !

Black holes are heavier than Regge states:  $M_{b.h.} = \frac{M_{\text{string}}}{g_{\text{string}}}$

Question: do loop and non-perturbative corrections change tree level signatures? Onset of n.p. physics:  $M_{b.h.}$ .

# Conclusions

- There exists many ISB models with SM like spectra without chiral exotics
- One can also have  
(Independent) String

INTERESTING TIMES FOR STRING  
PHENOMENOLOGY ARE AHEAD OF US.

THANK YOU !!

Computer  
Black

tions:  
metry!  
ons  
??  
 $M_{\text{string}}$   
 $g_{\text{string}}$

Question: do loop and non-perturbative corrections change tree level signatures? Onset of n.p. physics:  $M_{b.h.}$

Any null-vector  $k_i^2 = 0$  can be written in terms of two spinors  $(\lambda, \tilde{\lambda})$

$$\text{Momentum } k_i^\mu \longrightarrow \text{Dirac spinor} \begin{pmatrix} u_+(k_i)_\alpha \\ u_-(k_i)_{\dot{\alpha}} \end{pmatrix} \equiv \begin{pmatrix} (\lambda_i)_\alpha \\ (\tilde{\lambda}_i)_{\dot{\alpha}} \end{pmatrix}$$

$u(k)$  = Dirac spinor, helicity states  $u_\pm(k) = (1 \pm \gamma_5) u(k)$

with choice

$$u_+(k) = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{k^+} e^{i\varphi} \\ \sqrt{k^-} e^{i\varphi} \\ \sqrt{k^+} e^{i\varphi} \\ \sqrt{k^-} e^{i\varphi} \end{pmatrix}, \quad u_-(k) = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{k^-} e^{-i\varphi} \\ -\sqrt{k^+} \\ \sqrt{k^-} e^{-i\varphi} \\ \sqrt{k^+} \end{pmatrix} \quad k^\pm = k^0 \pm k^3$$

$$e^{\pm i\varphi} = \frac{k^1 \pm ik^2}{\sqrt{k^+ k^-}}$$

Define  $|i^\pm\rangle = u_\pm(k_i) \quad , \quad \langle i^\pm | = \overline{u_\pm(k_i)}$

Spinor products:

$$\langle ij \rangle := \langle i^- | j^+ \rangle = \overline{u_-(k_i)} u_+(k_j) \equiv \epsilon^{\alpha\beta} (\lambda_i)_\alpha (\lambda_j)_\beta = \sqrt{k_i k_j} e^{i\phi_{ij}},$$

$$[ij] := \langle i^+ | j^- \rangle = \overline{u_+(k_i)} u_-(k_j) \equiv \epsilon^{\dot{\alpha}\dot{\beta}} (\tilde{\lambda}_i)_{\dot{\alpha}} (\tilde{\lambda}_j)_{\dot{\beta}} = -\sqrt{k_i k_j} e^{-i\phi_{ij}}$$



$$\langle ij \rangle [ji] = -k_i k_j$$